

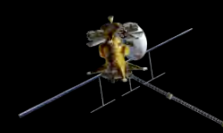
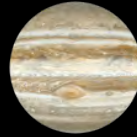
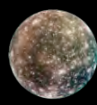
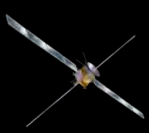
# Satellite Science from the Jupiter Europa Orbiter: An Element of the Europa Jupiter System Mission



Robert Pappalardo

Jet Propulsion Laboratory,  
California Institute of Technology

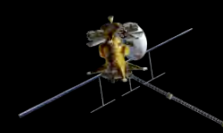
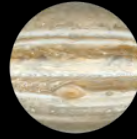
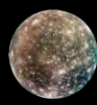
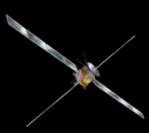
*A Joint NASA-ESA Outer Planet Mission Study*



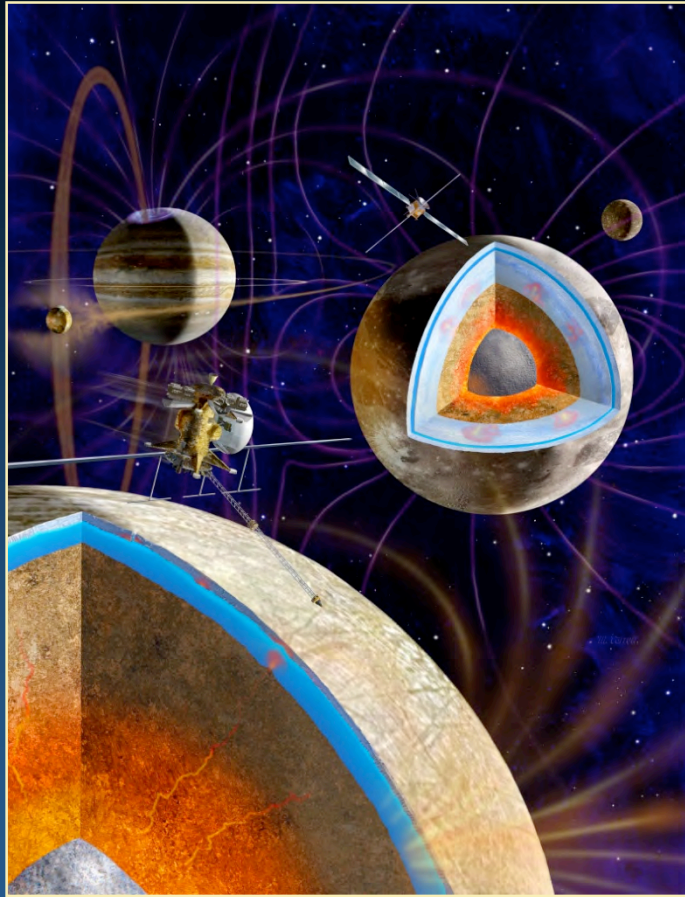
## 2002 Decadal Survey: Galilean Satellite Science

- Steering Group's "Science Themes" and associated "Fundamental Science Questions" point to Europa:
  - The Origin and Evolution of Habitable Worlds:
    - 7. How do planetary processes give rise to habitable zones and worlds?
  - Processes: How Planetary Systems Work:
    - 11. How do processes that shape planetary bodies operate and interact?
- Large Satellites Panel recommendations:
  - Robust Europa Orbiter with Jupiter system science is strongly recommended
  - Addresses 3 of 4 Large Satellites Panel panel themes and associated key questions
  - Next decade: Ganymede Orbiter, Io Explorer

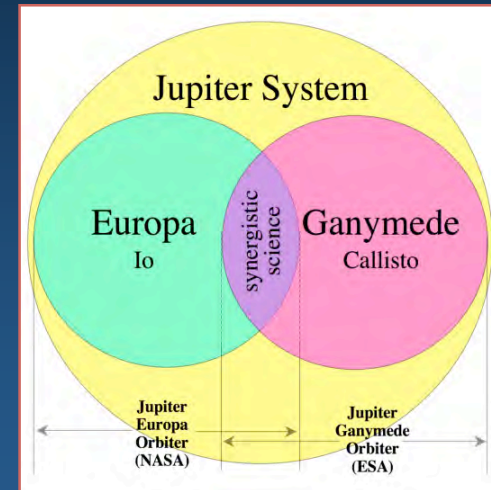
*2002 Decadal Survey recommended Europa as top flagship priority*



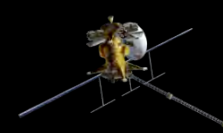
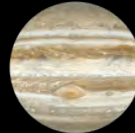
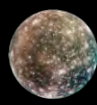
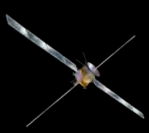
# Europa Jupiter Science Mission (EJSM)



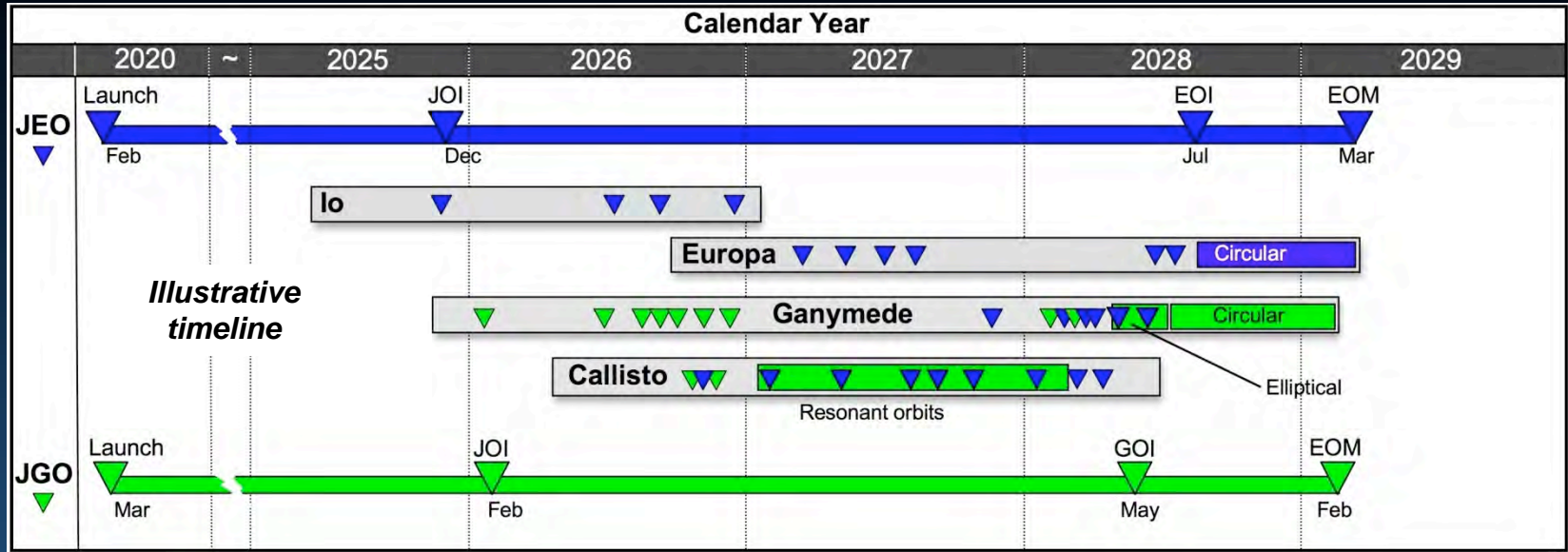
- NASA and ESA: Shared mission leadership
- Independently launched and operated orbiters
  - NASA-led Jupiter Europa Orbiter (JEO)
  - ESA-led Jupiter Ganymede Orbiter (JGO)
- Complementary science and payloads
  - JEO concentrates on Europa and Io
  - JGO concentrates on Ganymede and Callisto
  - Synergistic overlap
  - 11-12 instruments each
- Science goals:
  - Icy world habitability
  - Jupiter system processes



*Synergistic science: The sum of JEO + JGO is greater than the parts*

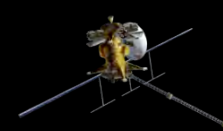
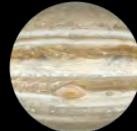
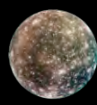
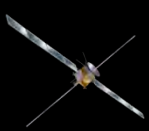


# Nominal EJSM Timeline



- Launches: 2020
- Jovian system tour phases: 2–3 years
- Moon orbital phases: 6–12 months
- End of Prime Missions: 2029
- Flexibility if either flight element is delayed or advanced

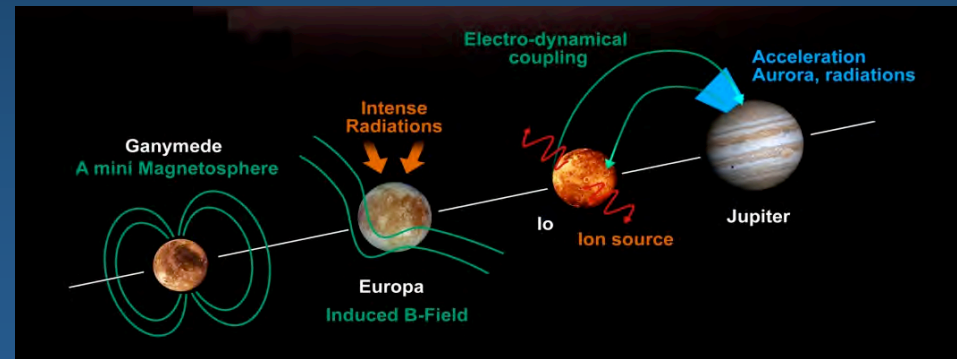
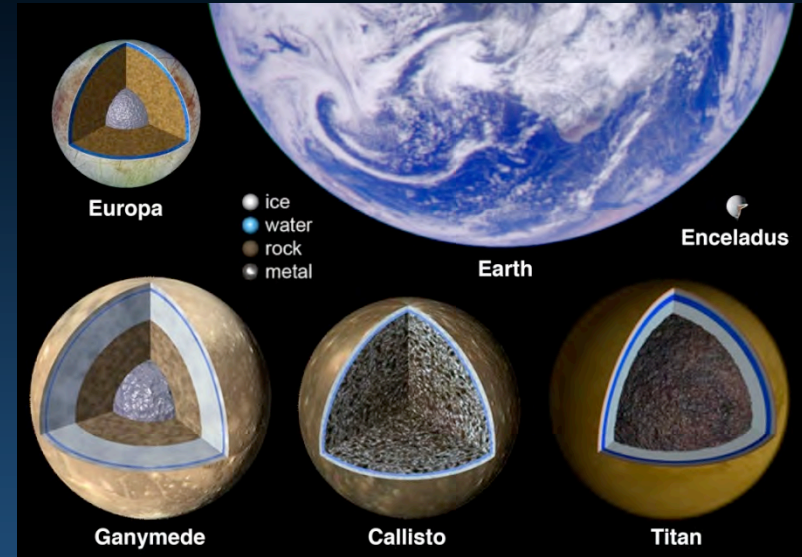
*Coordinated timelines ensure synergistic science*



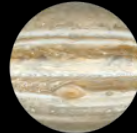
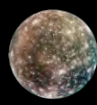
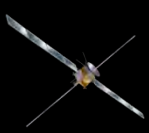
# EJSM Theme:

## The Emergence of Habitable Worlds Around Gas Giants

- **Goal 1: Determine if the Jupiter system harbors habitable worlds**
  - Ocean characteristics
  - Ice shells and subsurface water
  - Deep internal structure, and (for Ganymede) intrinsic magnetic field
  - External environments
  - Global surface compositions
  - Surface features and future landing sites
- **Goal 2: Characterize Jupiter system processes**
  - Satellite system
  - Jupiter atmosphere
  - Magnetodisk/magnetosphere
  - Jovian system Interactions
  - Jovian system origin

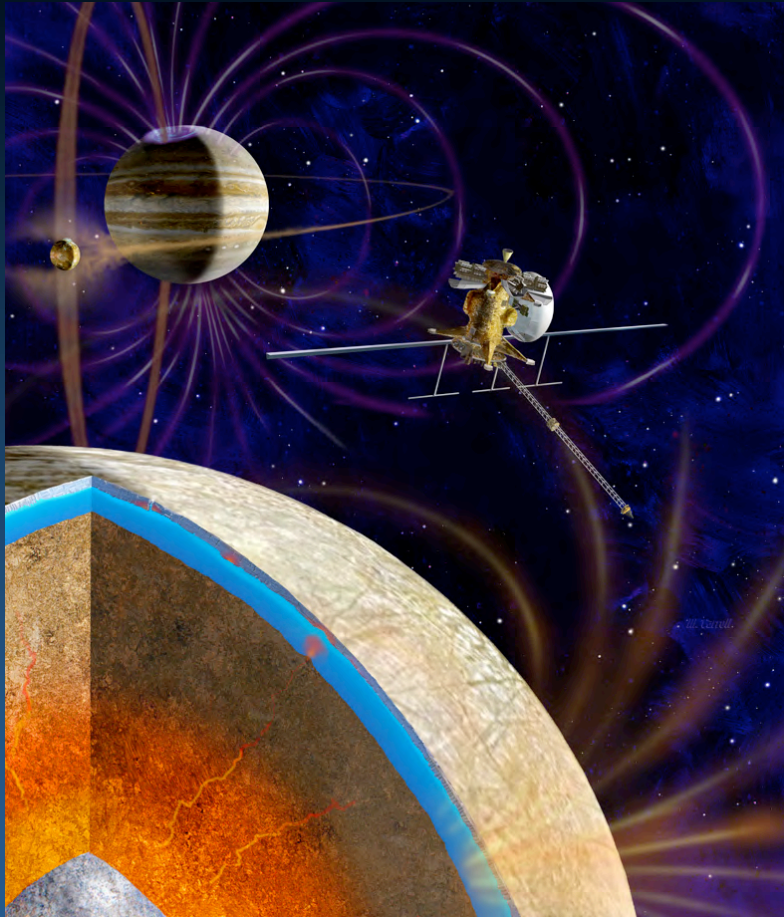


*Emphasis on icy moon habitability and Jupiter system processes*



# JEO Goal:

## Explore Europa to Investigate Its Habitability

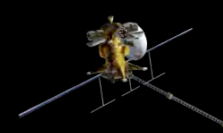
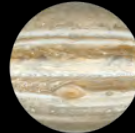
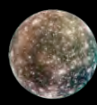
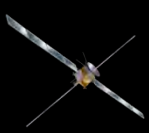


### Habitability

#### *Objectives (prioritized):*

- Ocean and Interior
- Ice Shell
- Chemistry and Composition
- Geology and Landing Sites
- Jupiter System
  - Satellite surfaces and interiors
  - Satellite atmospheres
  - Plasma and magnetospheres
  - Jupiter atmosphere
  - Rings

*Characterizing the archetype of icy world habitability*



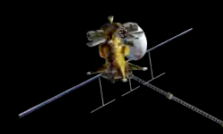
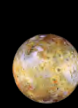
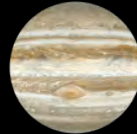
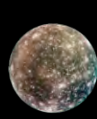
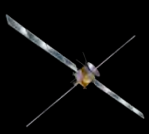
# JEO Traceability: Europa

Goal	Science Objective	Science Investigation
Explore Europa to investigate its habitability.	<b>A. Ocean</b> Characterize the extent of the ocean and its relation to the deeper interior.	A1. Determine the amplitude and phase of the gravitational tides.
		A2. Characterize the magnetic environment (including plasma) to determine the induction response from the ocean over multiple frequencies.
		A3. Characterize surface motion over the tidal cycle.
		A4. Determine the satellite's dynamical rotation state.
		A5. Investigate the core, rocky mantle, and rock-ocean interface.
	<b>B. Ice</b> Characterize the ice shell and any subsurface water, including their heterogeneity, and the nature of surface-ice-ocean exchange.	B1. Characterize the distribution of any shallow subsurface water.
		B2. Search for an ice-ocean interface.
		B3. Correlate surface features and subsurface structure to investigate processes governing material exchange among the surface, ice shell, and ocean.
		B4. Characterize regional and global heat flow variations.
	<b>C. Chemistry</b> Determine global surface compositions and chemistry, especially as related to habitability.	C1. Characterize surface organic and inorganic chemistry, including abundances and distributions of materials, with emphasis on indicators of habitability and potential biosignatures.
		C2. Relate compositions to geological processes, especially material exchange with the interior.
		C3. Characterize the global radiation environment and the effects of radiation on surface composition, atmospheric composition, albedo, sputtering, sublimation, and redox chemistry.
		C4. Characterize the nature of exogenic materials.
	<b>D. Geology</b> Understand the formation of surface features, including sites of recent or current activity, and identify and characterize candidate sites for future <i>in situ</i> exploration.	D1. Determine the formation history and three-dimensional characteristics of magmatic, tectonic, and impact landforms.
		D2. Determine sites of most recent geological activity, and evaluate future landing sites.
		D3. Investigate processes of erosion and deposition and their effects on the physical properties of the surface debris.

JEO Themes:

Origins Evolution Processes **Habitability** Life

Based on 2002 Decadal's "objectives of solar system exploration"

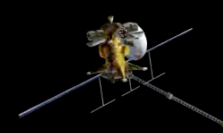
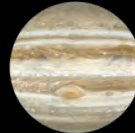
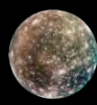
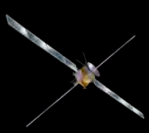


# JEO Traceability: Jupiter System Science

Goal	Science Objective	Science Investigation	
Explore Europa to investigate its habitability.	Understand Europa in the context of the Jupiter system.	Satellite surfaces and interiors	E1. Investigate the nature and magnitude of tidal dissipation and heat loss on the Galilean satellites, particularly Io
			E2. Investigate Io's active volcanism for insight into its geological history and evolution (particularly of its silicate crust)
			E3. Investigate the presence and location of water within Ganymede and Callisto.
			E4. Determine the composition, physical characteristics, distribution and evolution of surface materials on Ganymede.
			E5. Determine the composition, physical characteristics, distribution and evolution of surface materials on Callisto.
			E6. Identify the dynamical processes that cause internal evolution and near-surface tectonics of Ganymede and Callisto.
		Satellite Atms.	E7. Characterize the composition, variability and dynamics of Europa's atmosphere and ionosphere
			E8. Understand the sources and sinks of Io's crustal volatiles and atmosphere.
			E9. Determine the sources and sinks of the Ganymede and Callisto atmospheres.
		Plasma and magnetospheres	E10. Characterize the neutral atoms and molecules escaping Europa's gravity.
			E11. Characterize the composition of and transport in Io's plasma torus.
			E12. Study the pickup and charge exchange processes in the Jupiter system plasma and neutral tori.
			E13. Study the interactions between Jupiter's magnetosphere and Io, Ganymede and Callisto (incl. characterize Ganymede's magnetic field)
			E14. Understand the structure, composition and stress balance of Jupiter's magnetosphere.
		Jupiter atm.	E15. Determine how plasma and magnetic flux are transported in Jupiter's magnetosphere.
			E16. Characterize the abundance of minor species (especially water and ammonia) in Jupiter's atmosphere to understand the evolution of the Jovian system, including Europa.
		Rings	E17. Characterize Jovian atmospheric dynamics and structure.
			E18. Characterize the properties of the small moons, ring source bodies and dust
			E19. Identify the dynamical processes that define the origin and dynamics of ring dust.

E. Jupiter system science

Giant Planets Panel

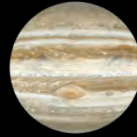
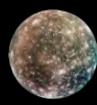
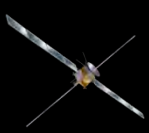


# JEO Model Payload

<b>Ocean</b>		
Laser Altimeter		LA
Radio Science		RS
<b>Ice</b>		
Ice Penetrating Radar		IPR
<b>Chemistry</b>		
Vis-IR Imaging Spectrometer		VIRIS
UV Spectrometer		UVS
Ion and Neutral Mass Spectrometer		INMS
<b>Geology</b>		
Thermal Instrument		TI
Narrow Angle Camera		NAC
Wide Angle Camera and Medium Angle Camera		WAC + MAC
<b>Fields and Particles</b>		
Magnetometer		MAG
Particle and Plasma Instrument		PPI

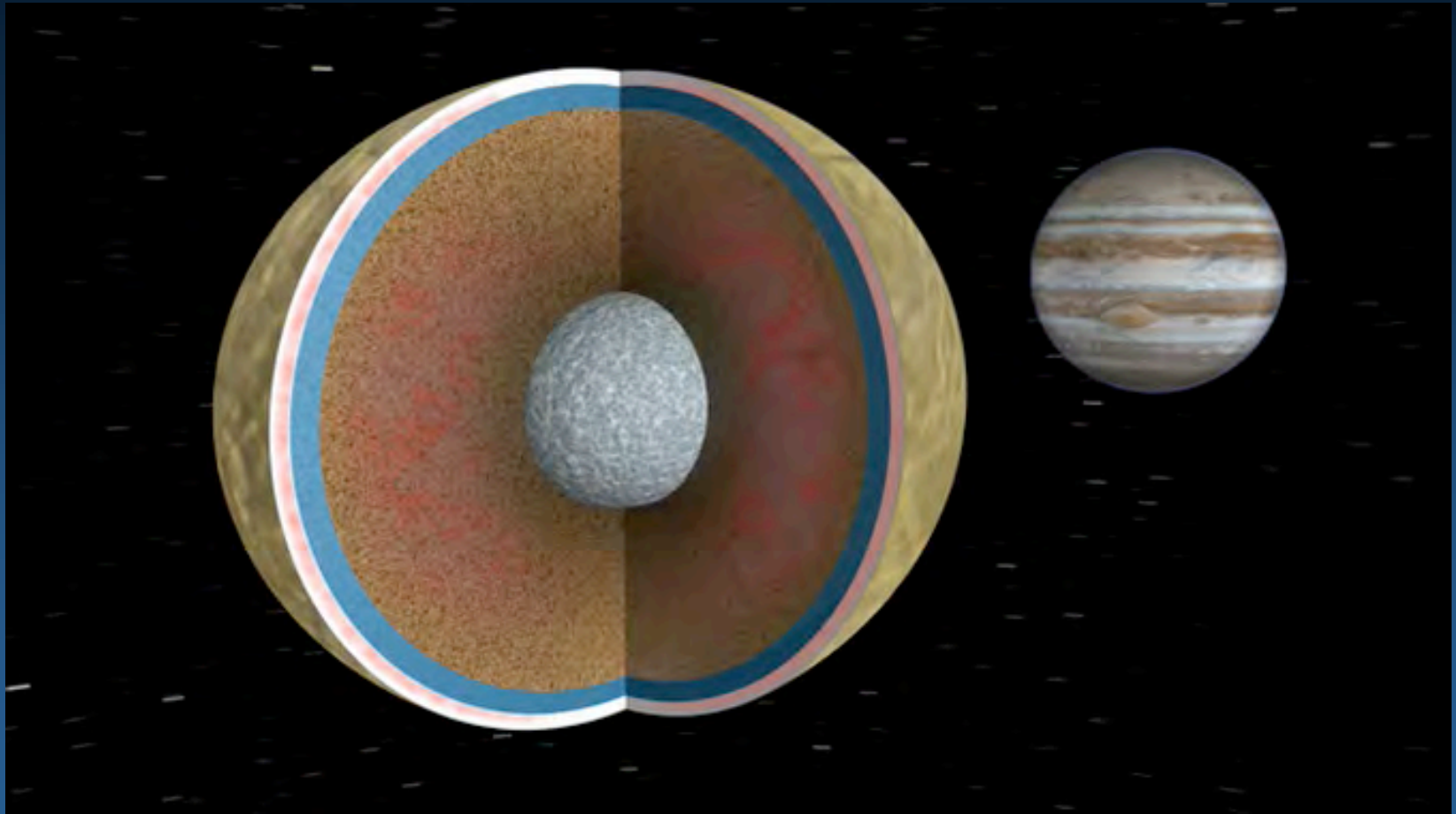
- Model payload is a proof-of-concept example
  - Other instrument choices may be viable
- Emphasizes accomplishing Europa investigations
- Enables robust Jupiter system science
- The final selected payload will almost certainly be different

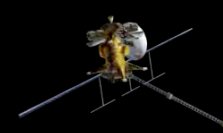
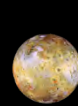
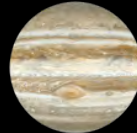
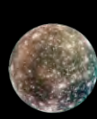
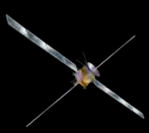
*Capable model payload with a conservative approach*



# Europa: Ocean • Ice • Chemistry • Geology

*A. Characterize the extent of the ocean and its relation to the deeper interior.*

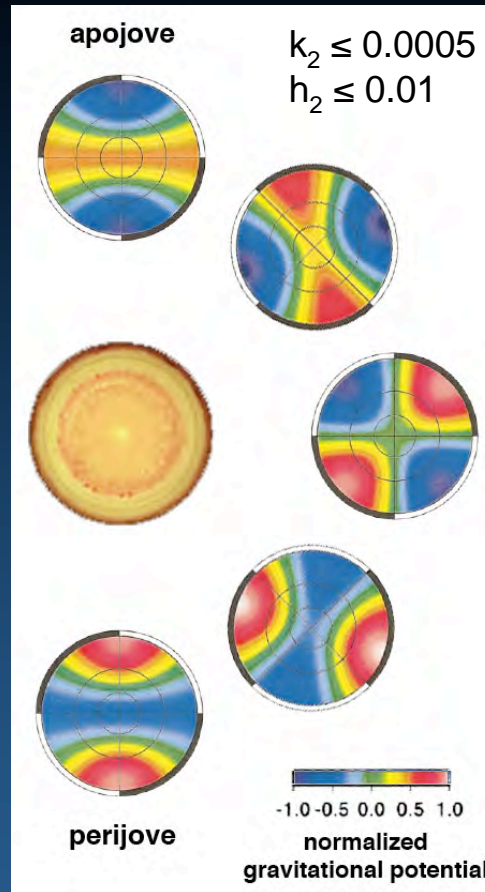




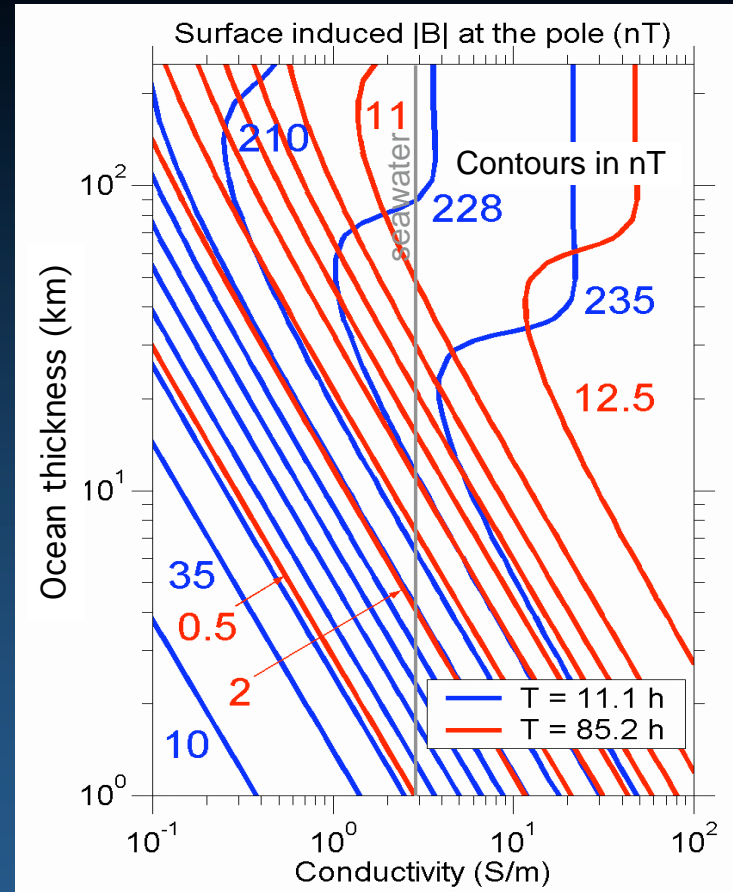
# Europa: Ocean • Ice • Chemistry • Geology

## A. Ocean & deeper interior:

1. Gravitational tides
2. Magnetic environment (including plasma)
3. Surface motion
4. Dynamical rotation state
5. Core, rocky mantle, & rock-ocean interface

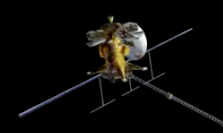
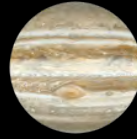
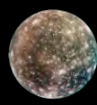
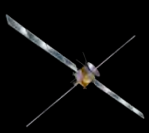


[Moore & Schubert, 2000]

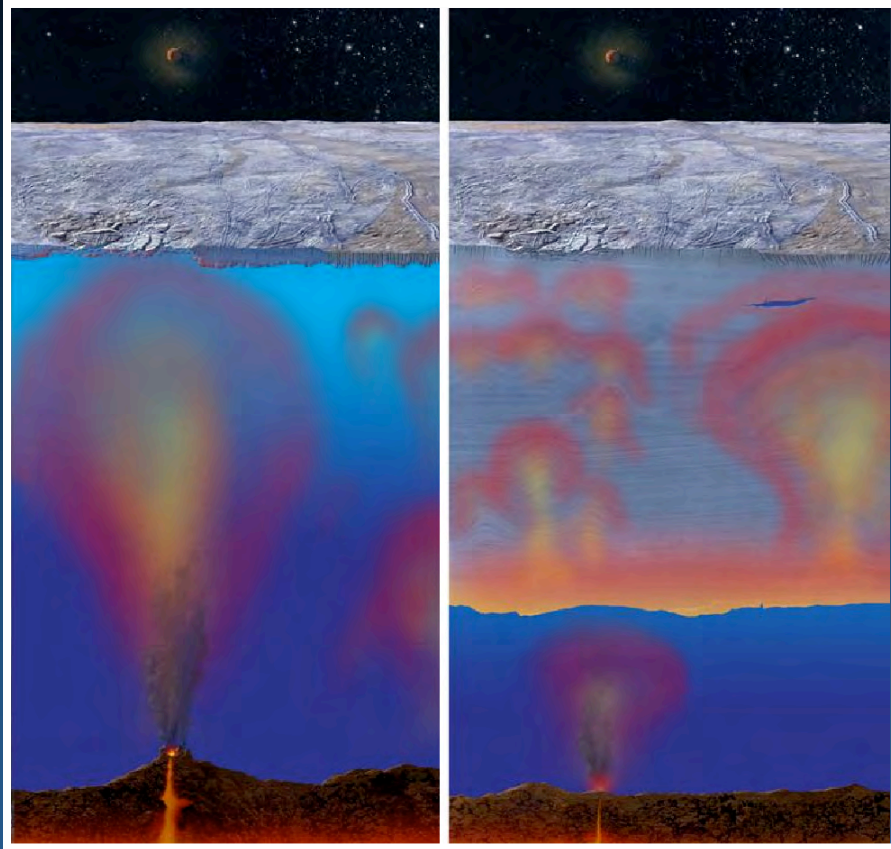


[Khurana, 2002]

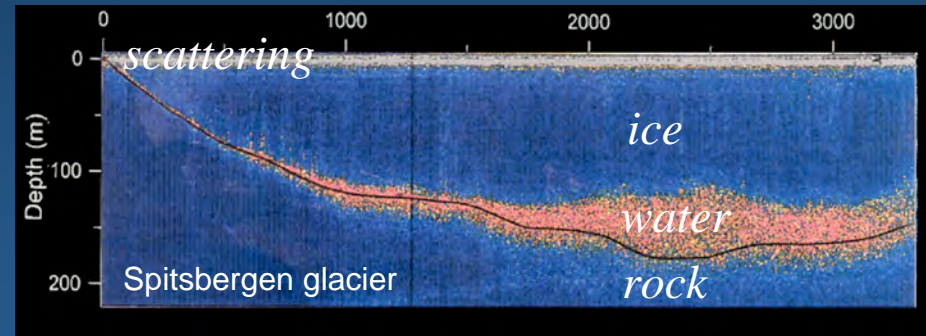
*Geophysical techniques reveal the nature of the interior*



# Europa: Ocean • Ice • Chemistry • Geology

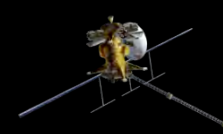
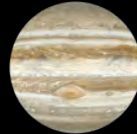
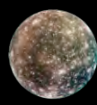
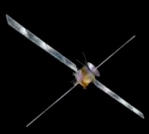


*B. Characterize the ice shell and any subsurface water, including their heterogeneity, and the nature of surface-ice-ocean exchange*



[Bjornsson et al., 1996]

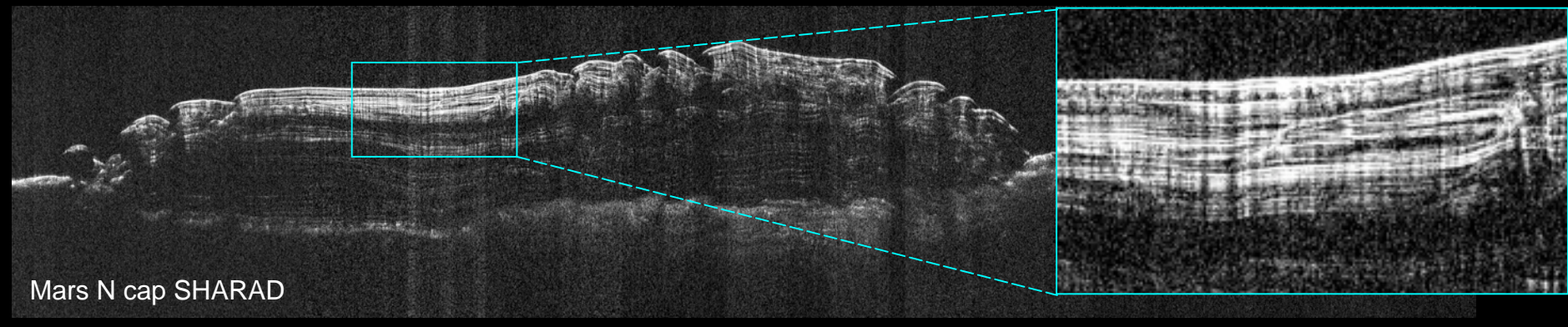
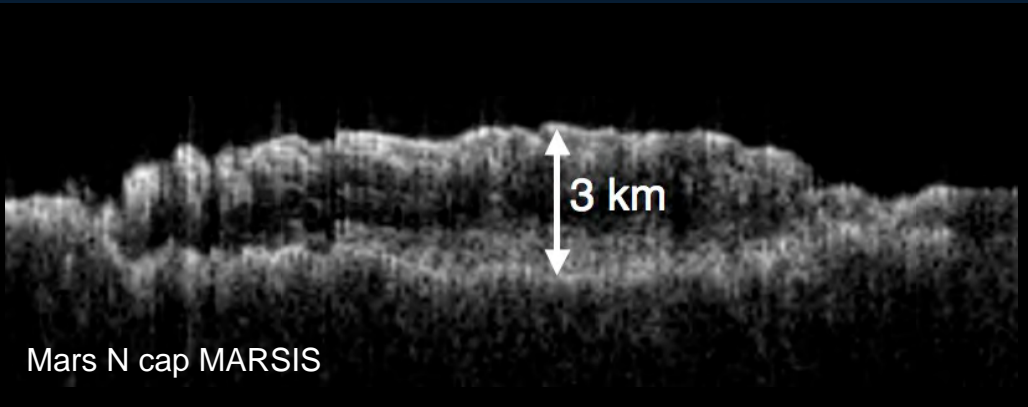
***Exchange processes are critical to Europa's habitability***



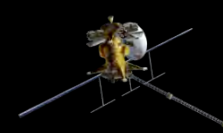
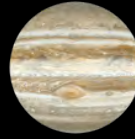
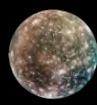
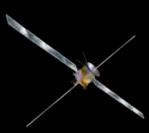
# Europa: Ocean • Ice • Chemistry • Geology

## *B. Ice shell & subsurface water:*

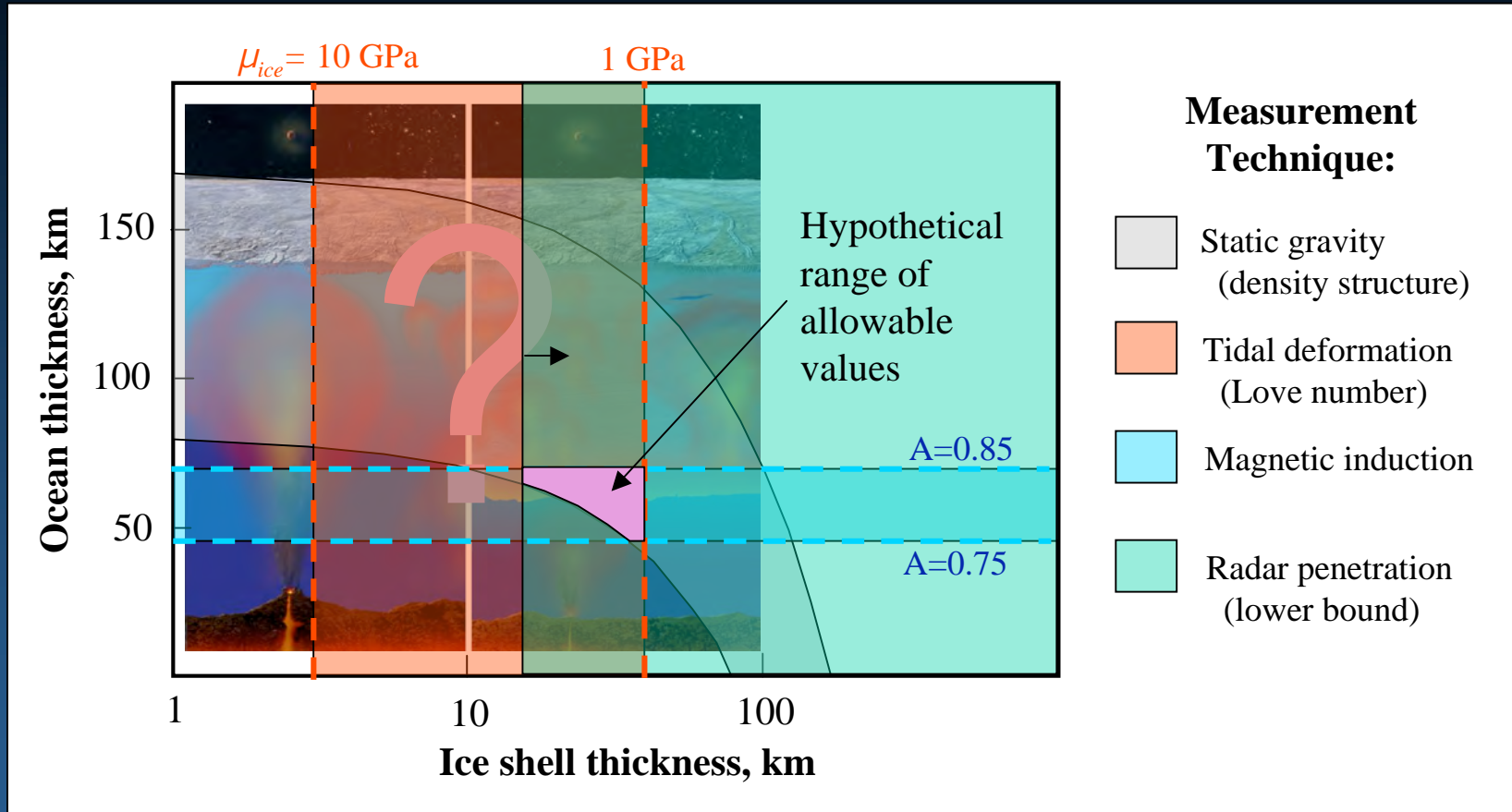
1. Shallow water
2. Ice-ocean interface
3. Material exchange
4. Heat flow variations



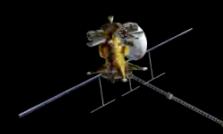
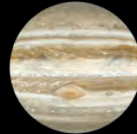
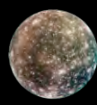
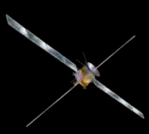
*Radar sounding would characterize the ice shell in 3 dimensions*



# Constraining Ice Shell Thickness: *Hypothetical Example*



*Multiple techniques constrain ice shell thickness*

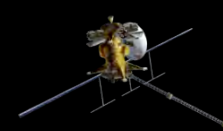
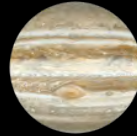
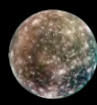
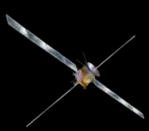


# Europa: Ocean • Ice • Chemistry • Geology

*C. Determine global surface compositions and chemistry, especially as related to habitability*



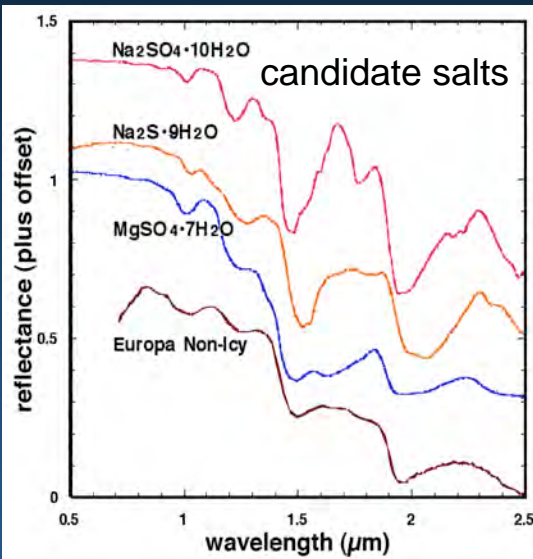
*Composition is key to understanding ocean habitability*



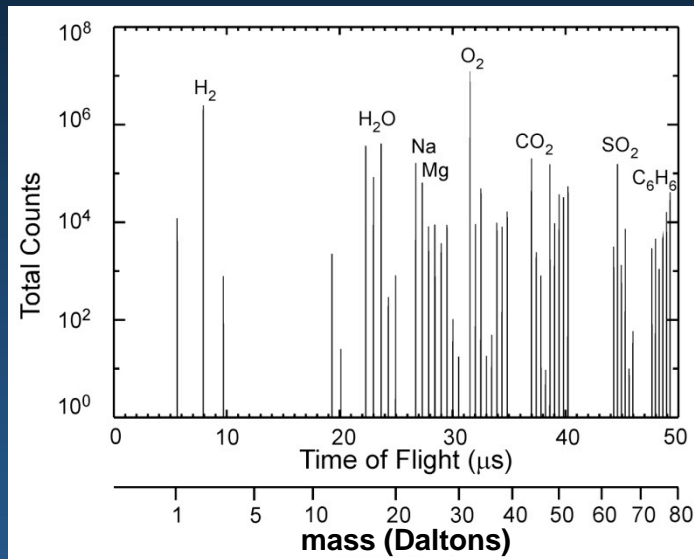
# Europa: Ocean • Ice • Chemistry • Geology

## *C. Global surface composition & chemistry:*

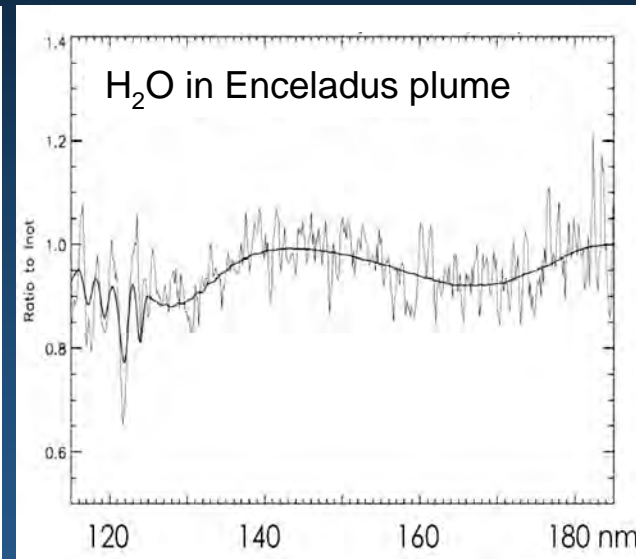
1. Organic & inorganic chemistry
2. Relation to geologic processes
3. Radiation effects
4. Exogenic materials



IR spectroscopy:  
surface constituents

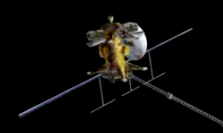
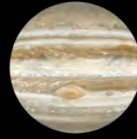
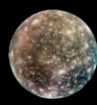
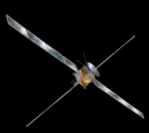


INMS:  
sputtered particles



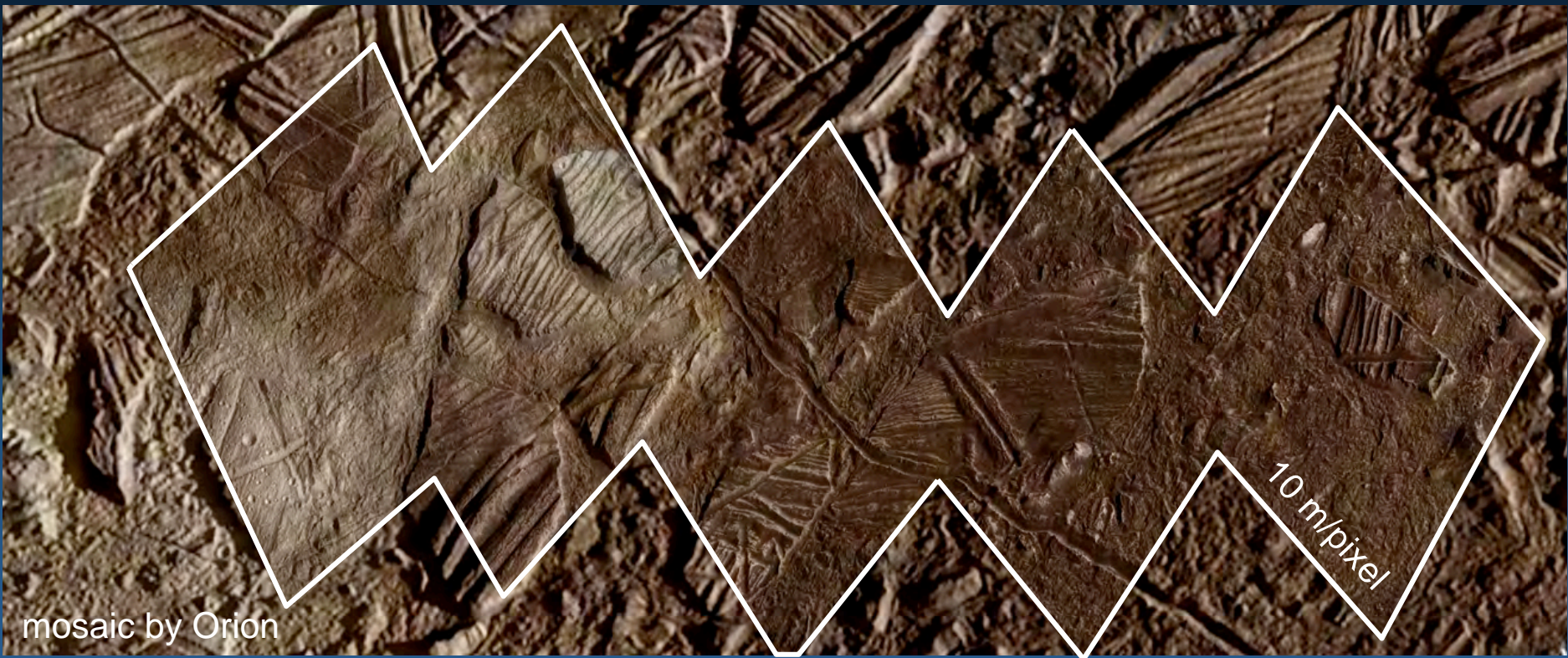
Stellar occultations:  
atmospheric species

*Multiple techniques characterize and map composition*



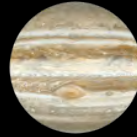
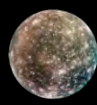
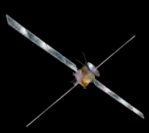
# Europa: Ocean • Ice • Chemistry • Geology

*D. Understand the formation of surface features, including sites of recent or current activity, and identify and characterize candidate sites for future in situ exploration*



mosaic by Orion

***JEO would increase 10-20 m/pixel image coverage from 0.03% to 7%***

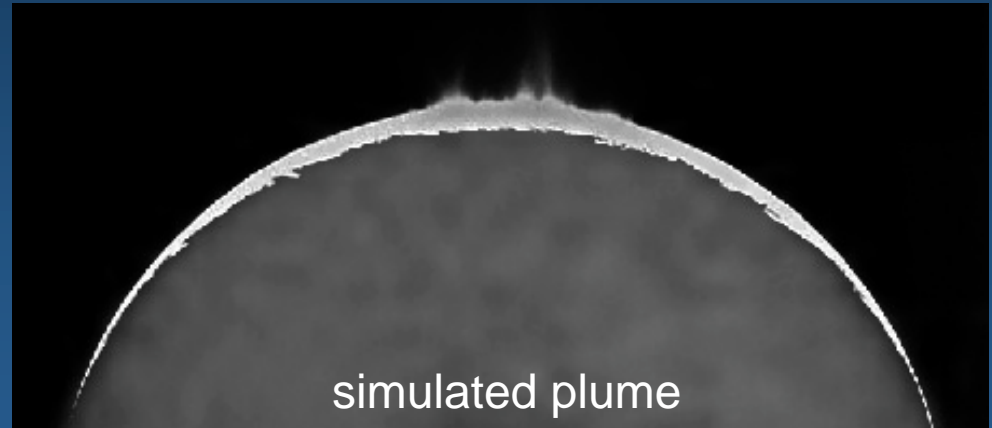


# Europa: Ocean • Ice • Chemistry • Geology

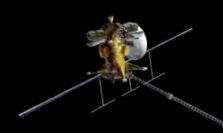
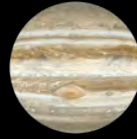
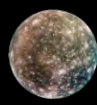
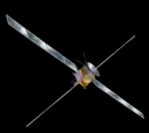


## *D. Surface features, activity, & landing sites:*

1. Formation history & 3-D character
2. Recent activity & potential future landing sites
3. Erosion & deposition

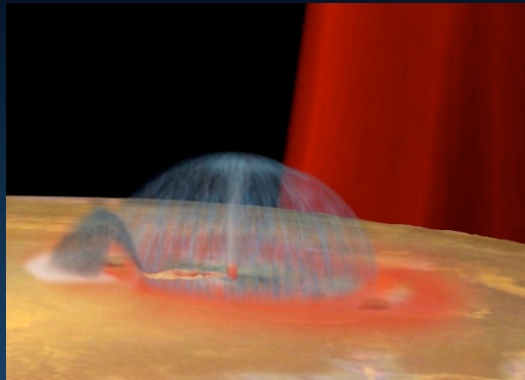


*JEO would decipher Europa's varied and complex geology*

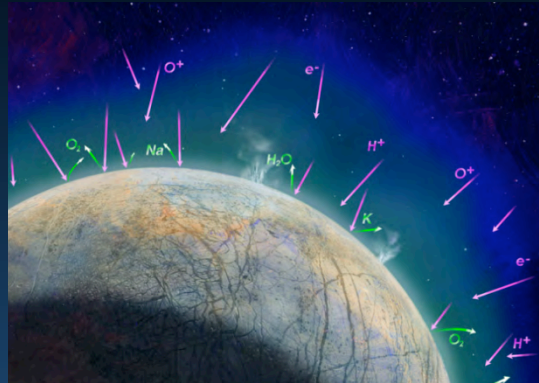


# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

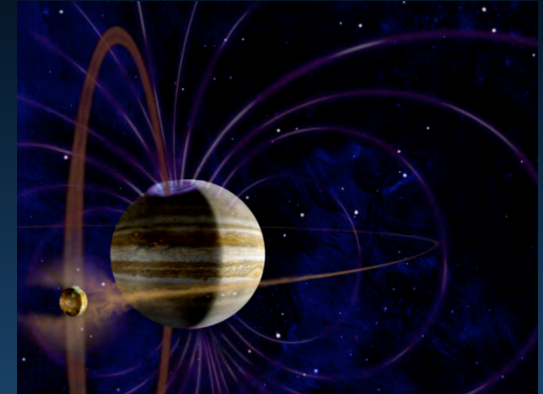
*E. Understand Europa in the context of the Jupiter system*



Satellite surfaces & interiors



Satellite atmospheres



Plasma & magnetospheres

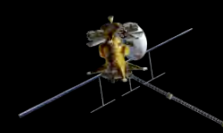
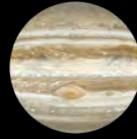
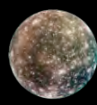
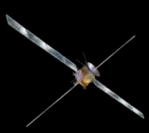


Jupiter atmosphere



Rings

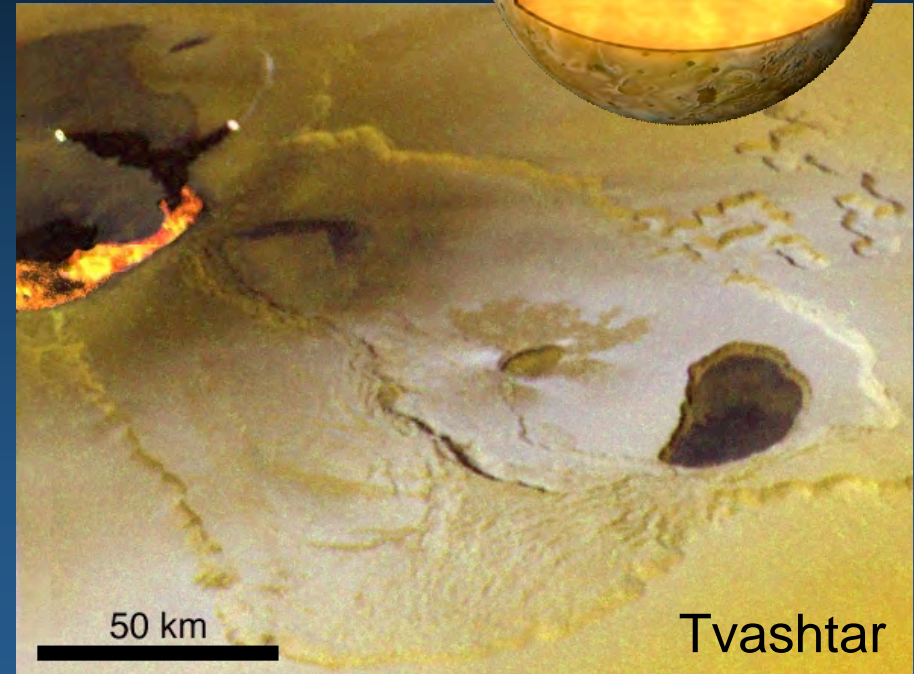
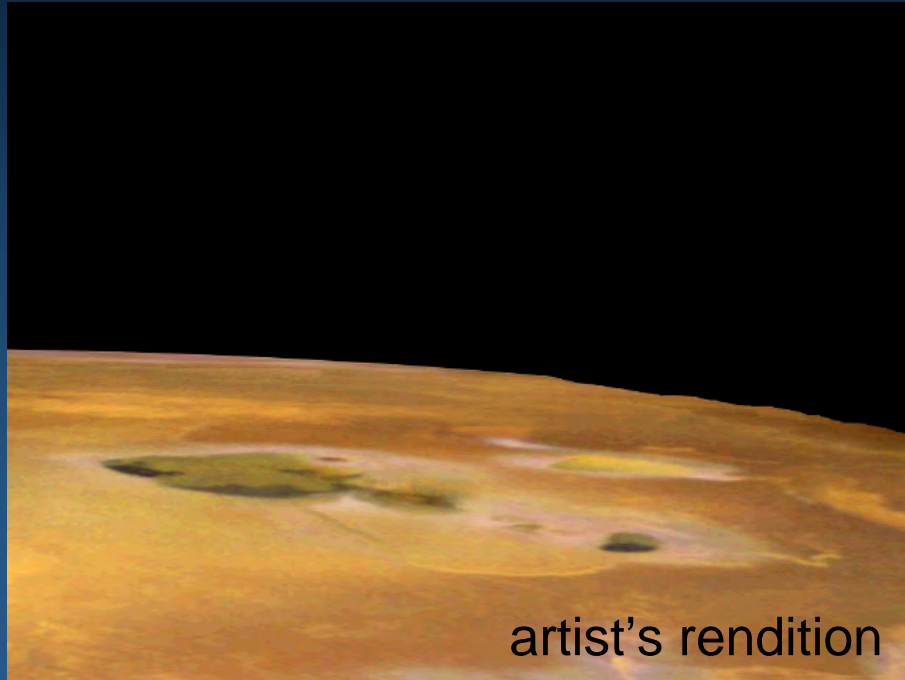
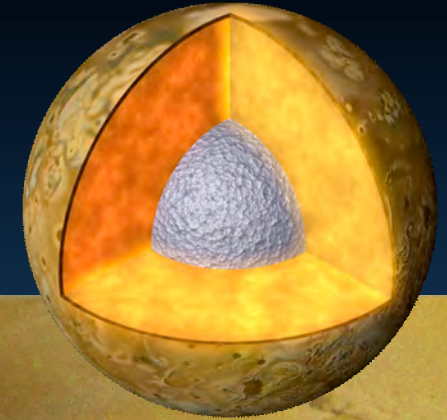
*The Jupiter system is rich in dynamic and coupled processes*



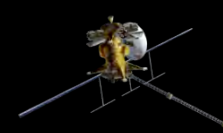
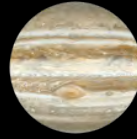
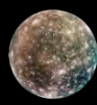
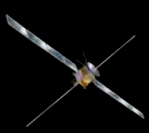
# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

## Satellite surfaces & interiors:

1. Io's tidal heating & heat loss
2. Io's active volcanism



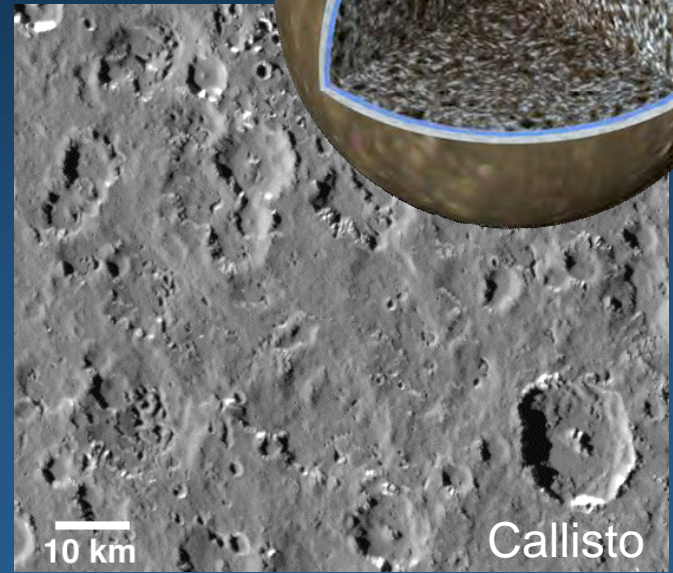
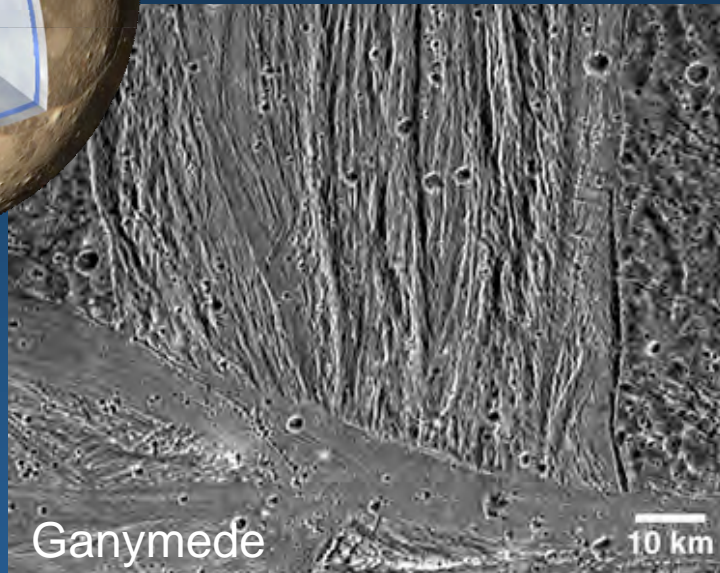
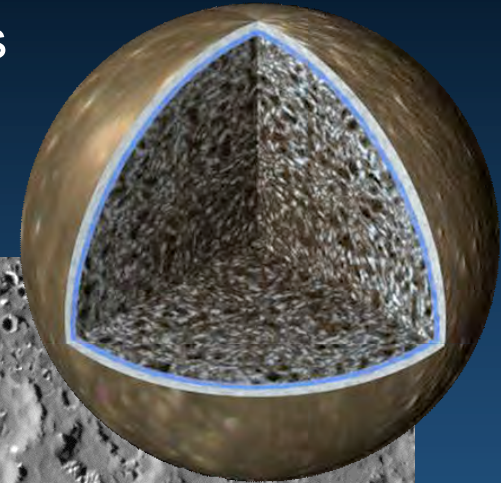
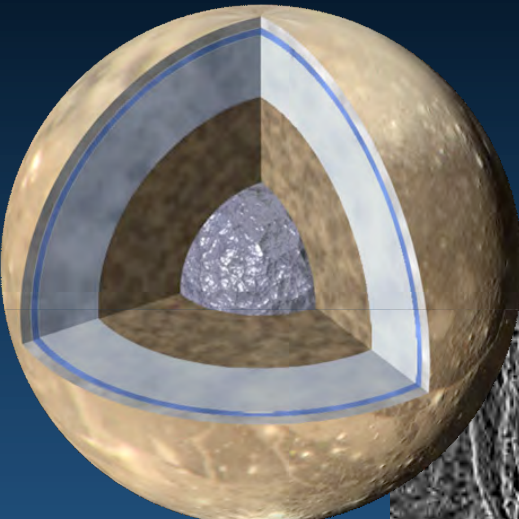
*Io is the tidal engine of the Laplace resonance*



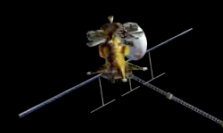
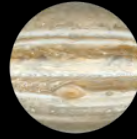
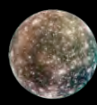
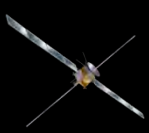
# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

## *Satellite surfaces & interiors (cont.):*

3. Water in Ganymede & Callisto
4. Ganymede's surface materials
5. Callisto's surface materials
6. Internal evolution & tectonics



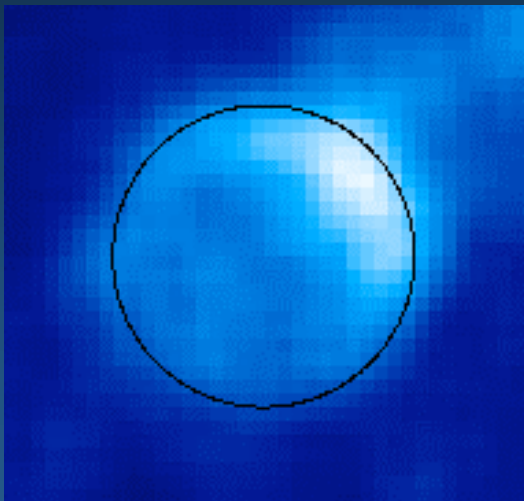
*The icy Galilean satellites provide context for Europa*



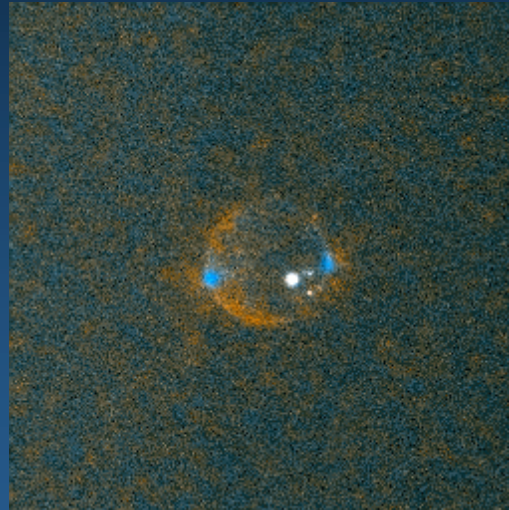
# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

## *Satellite atmospheres:*

7. Europa: Composition, variability, and dynamics
8. Io: Composition, sources, and evolution
9. Ganymede and Callisto: Sources and sinks



Europa atm. (HST O 1356Å)

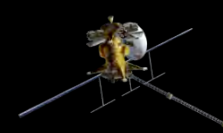
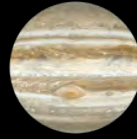
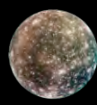
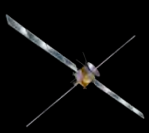


Io atm. & aurora (Cassini)

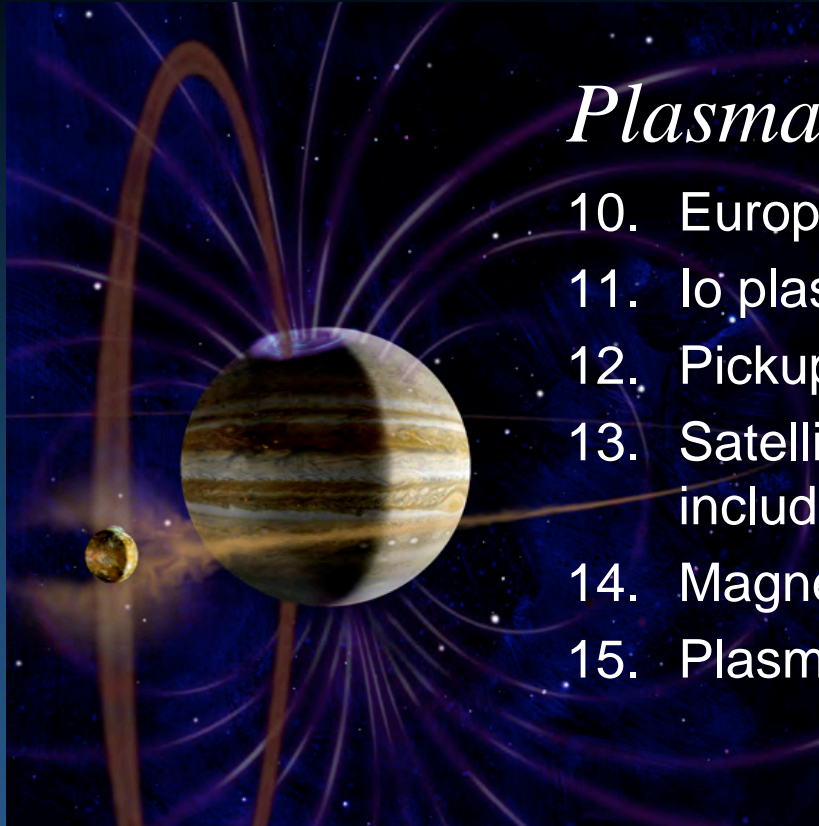


Ganymede aurora (HST)

*Understanding atmospheric interactions and processes*

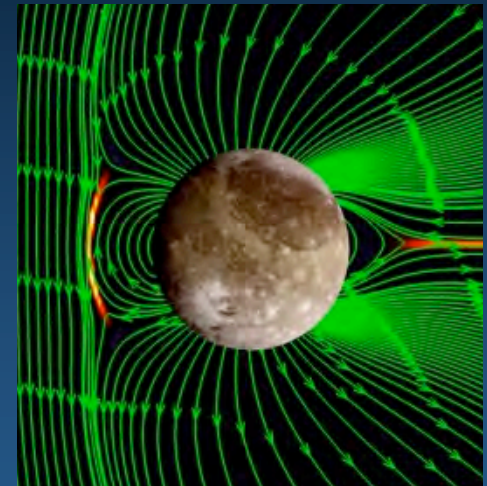


# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

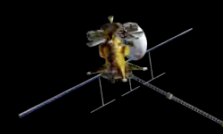
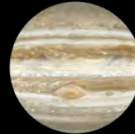
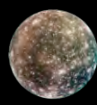
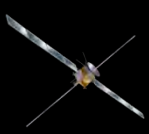


## *Plasma & magnetospheres:*

- 10. Europa's escaping neutrals
- 11. Io plasma torus
- 12. Pickup & charge exchange
- 13. Satellite interactions including Ganymede's field
- 14. Magnetospheric structure
- 15. Plasma transport



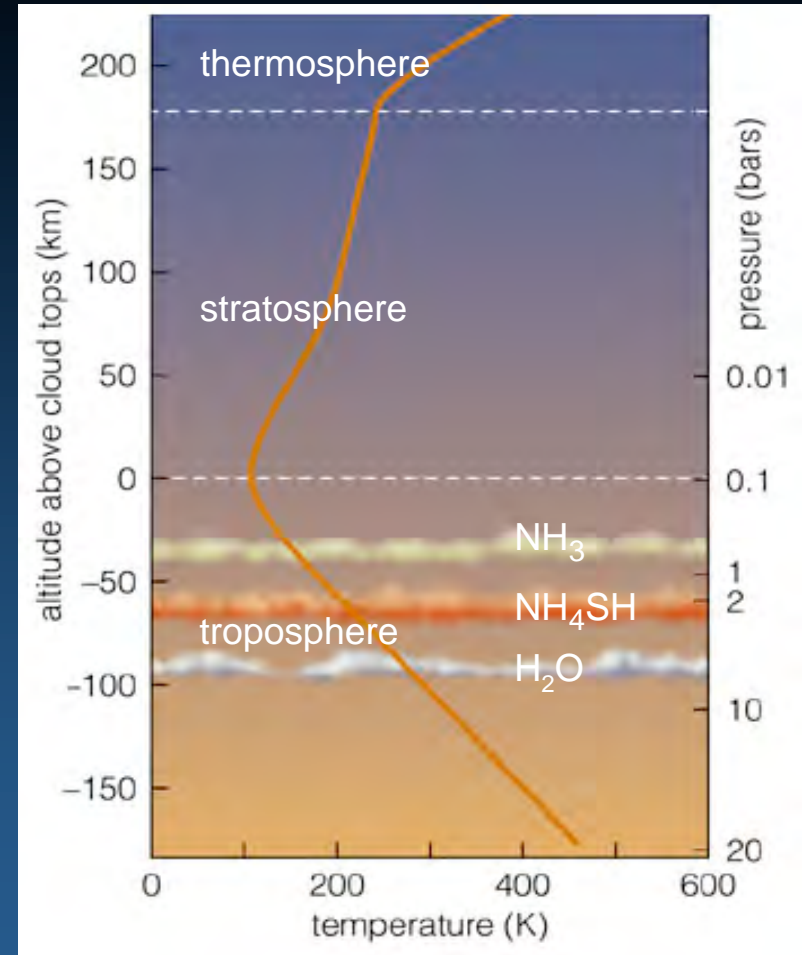
*Probing the Solar System's largest magnetosphere and its unique satellite-magnetosphere interactions*



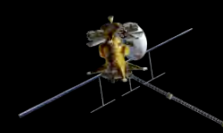
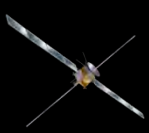
# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

## *E. Jupiter atmosphere:*

- 16. Minor species abundances
- 17. Dynamics & structure



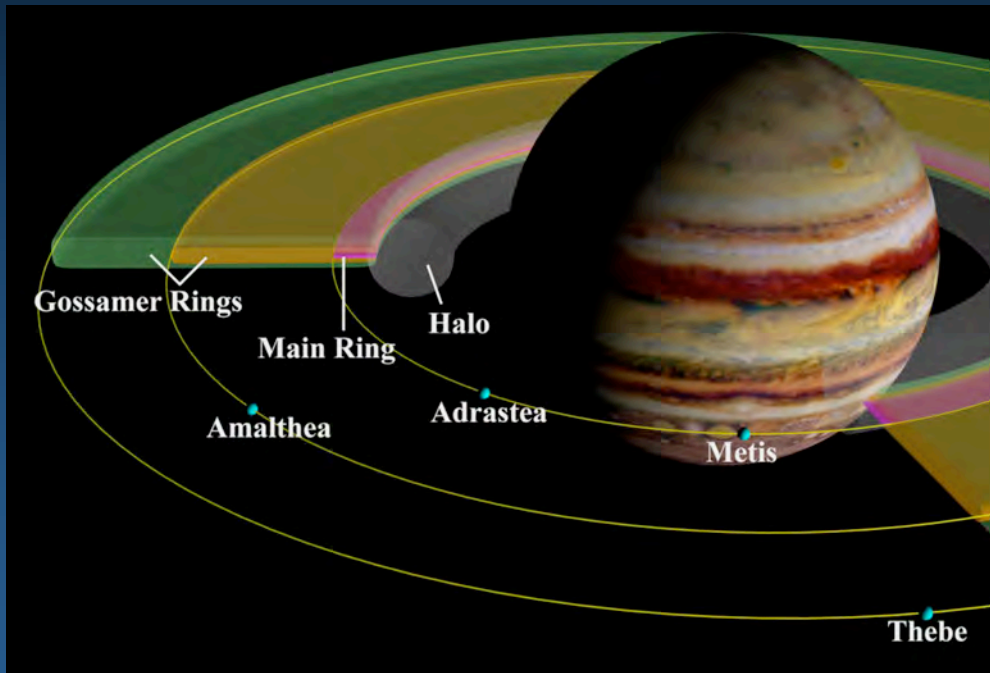
*Addresses unanswered questions and complements Juno*



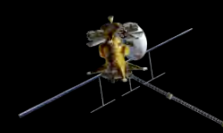
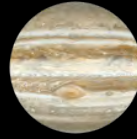
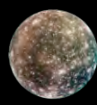
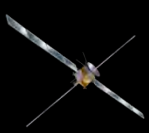
# Jupiter System: Sats • Atm • Mag • Jupiter • Rings

## *E. Rings:*

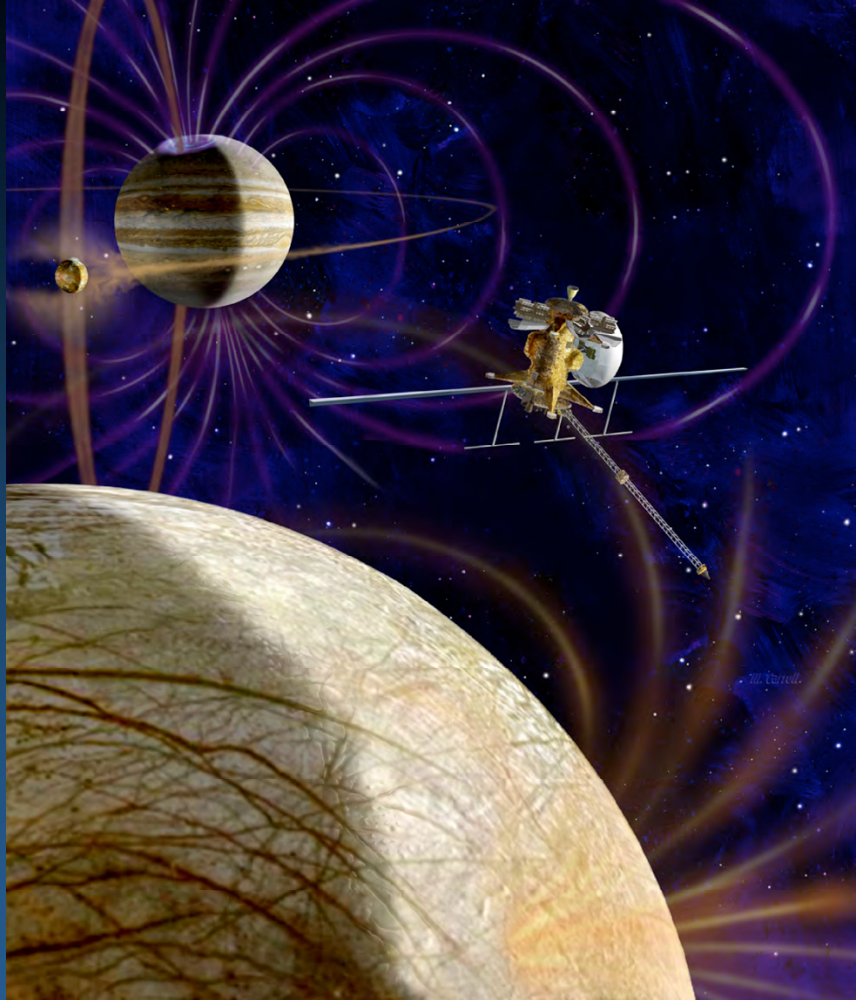
- 18. Ring source bodies
- 19. Dynamical processes



*Comparative studies of ring dynamics and evolution*



# Jupiter Europa Orbiter Science

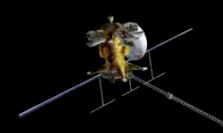
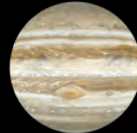
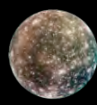
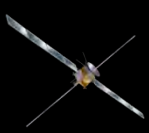


## Habitability

### *Objectives Summary:*

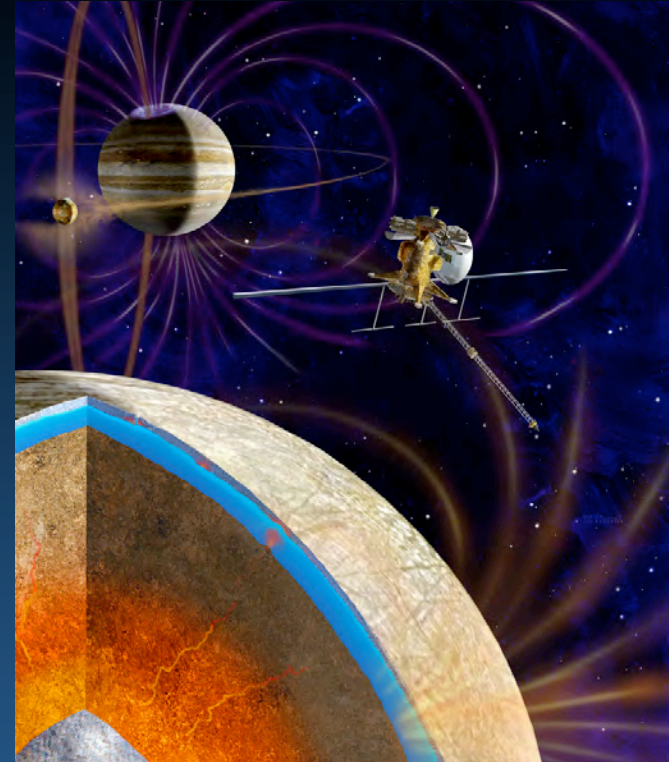
- Ocean characterization
- Surface-ice-ocean exchange
- Compositional makeup
- Geological evolution
- Jupiter system science
  - Galilean satellite evolution
  - Sat. atmospheric interactions
  - Magnetospheric physics
  - Jupiter atmospheric dynamics
  - Ring system dynamics

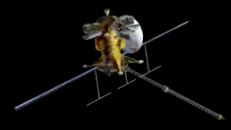
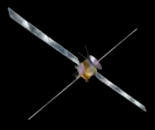
*Rich and robust science of Europa and the Jupiter system*



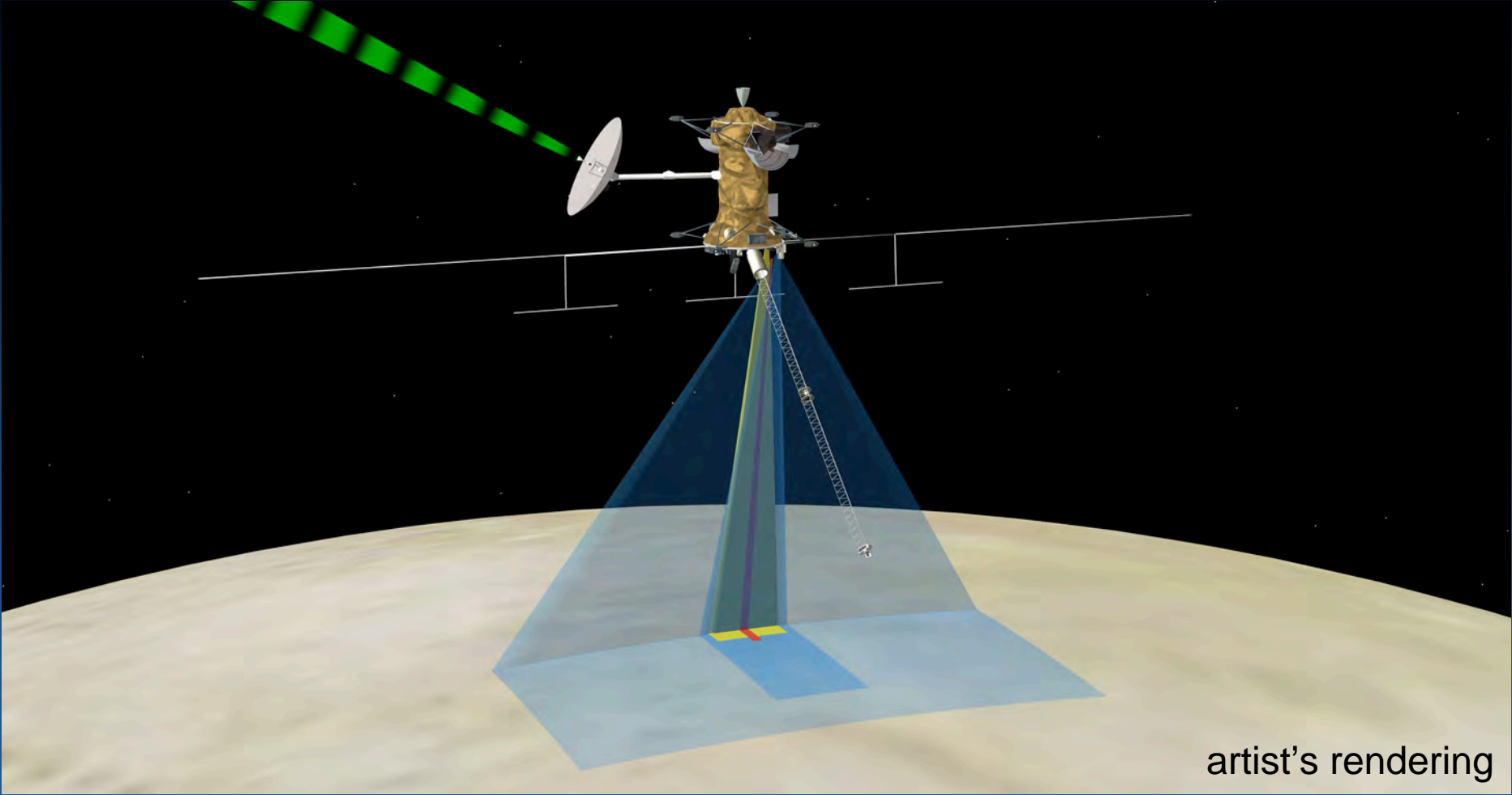
# JEO Baseline Mission Overview

- Launch vehicle: Atlas V 551
- Power source: 5 MMRTG or 5 ASRG
- Mission timeline:
  - Launch: 2020
  - Jovian system tour phase: 30 mo.
  - Europa orbital phase: 9 mo.
  - End of prime mission: 2029
  - Spacecraft final disposition is Europa impact
- Radiation dose: 2.9 Mrad (behind 100 mils Al)
  - Handled using a combination of rad-hard parts and tailored component shielding
  - Key rad-hard parts are available, with the required heritage
  - Team is developing and providing design information and approved parts list for prospective suppliers of components, including instruments



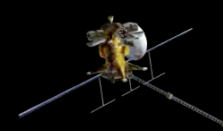
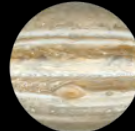
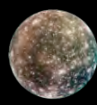
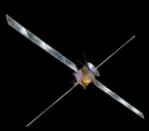


# Orbital Operability

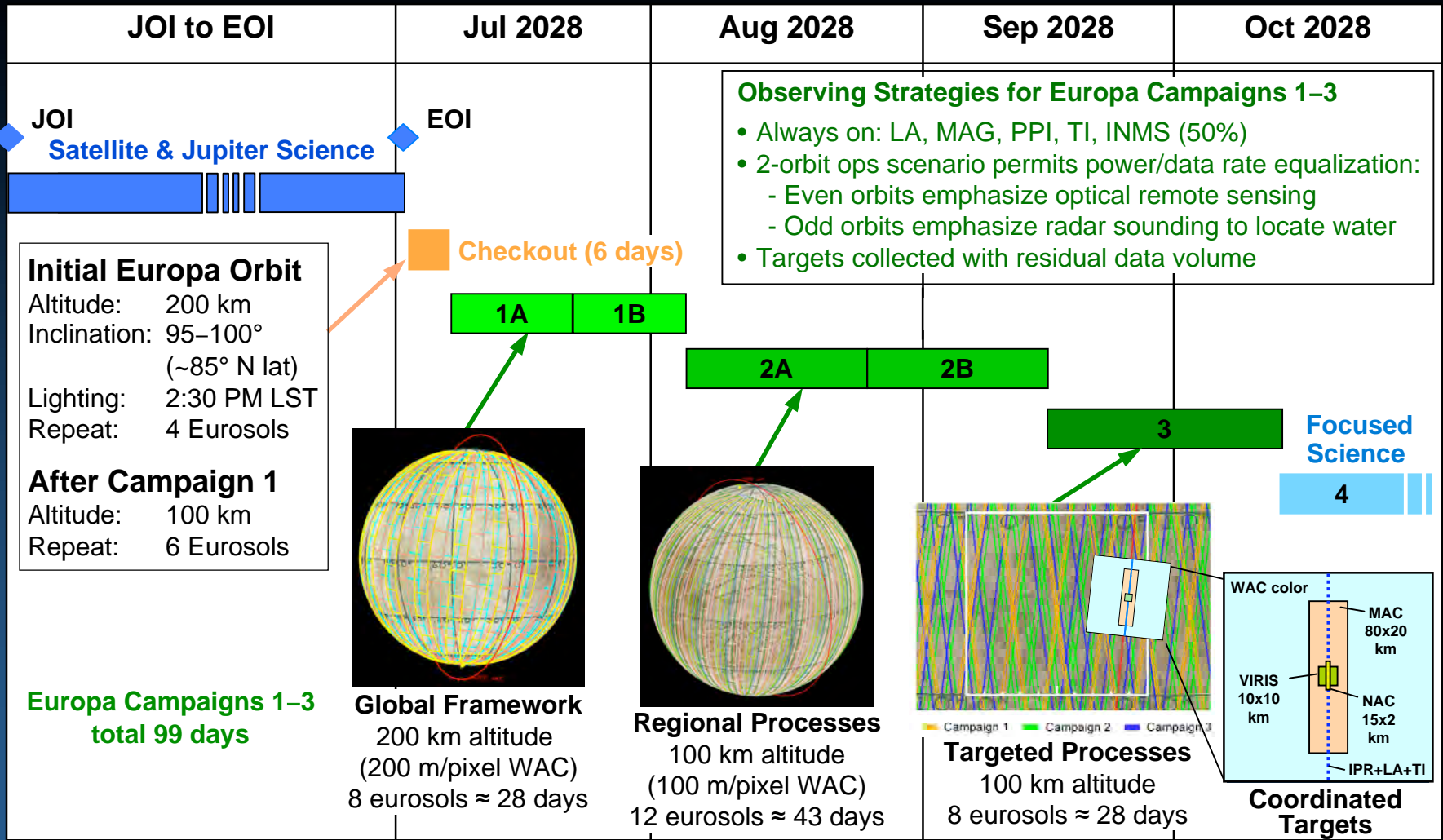


artist's rendering

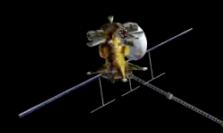
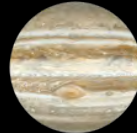
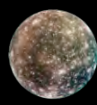
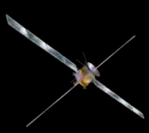
*Articulated antenna would permit simultaneous observations and downlink*



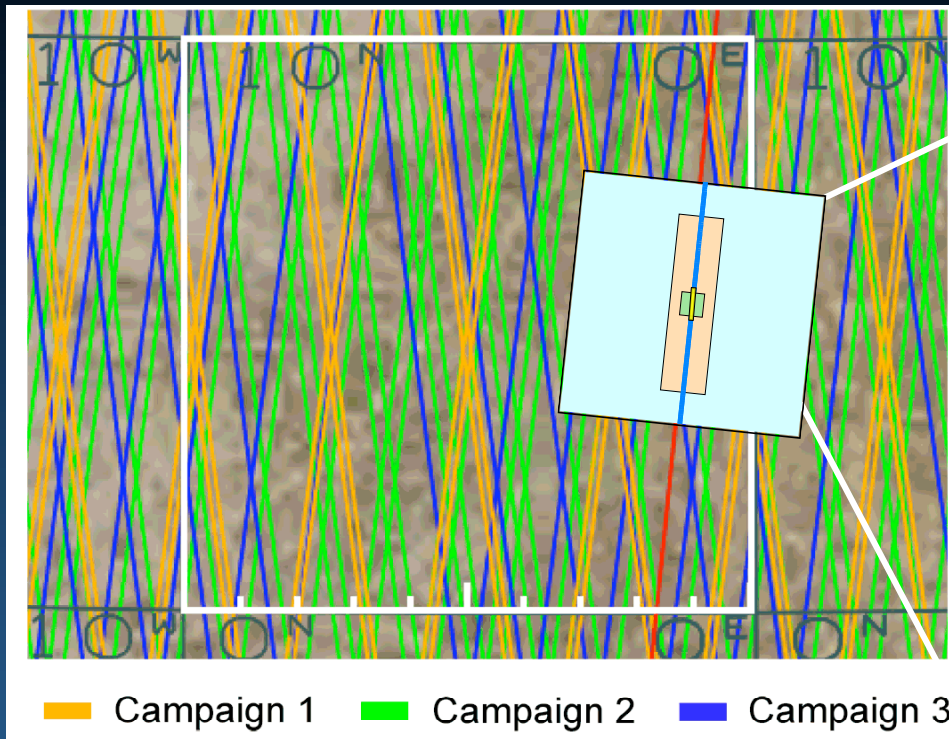
# Europa Science Campaigns



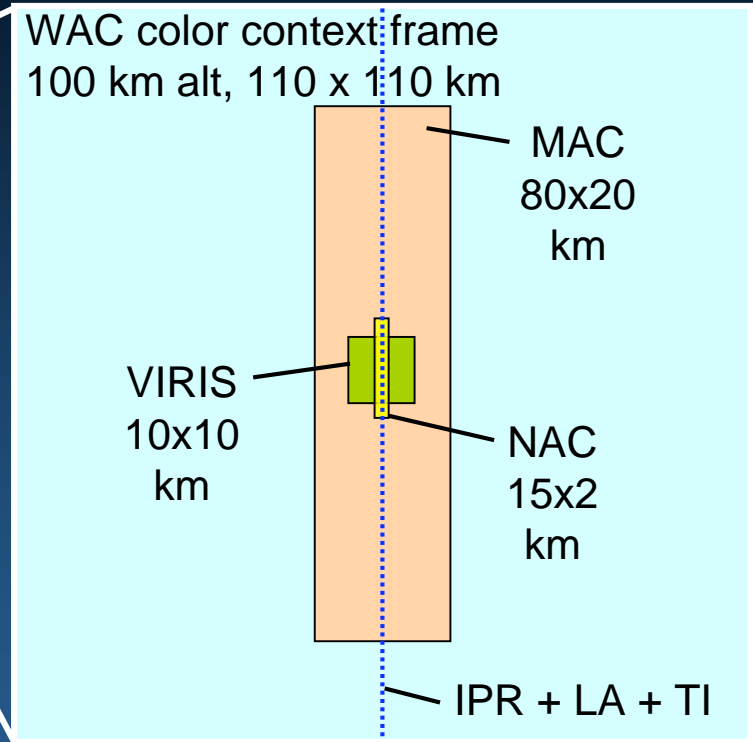
*Europa science objectives addressed in first 100 days in orbit*



# Europa Science Campaigns: Profiling and Targeted Observations

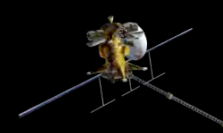
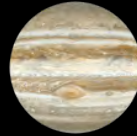
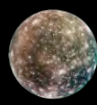
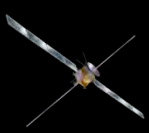


≤18 km groundtrack separation after 100 dy

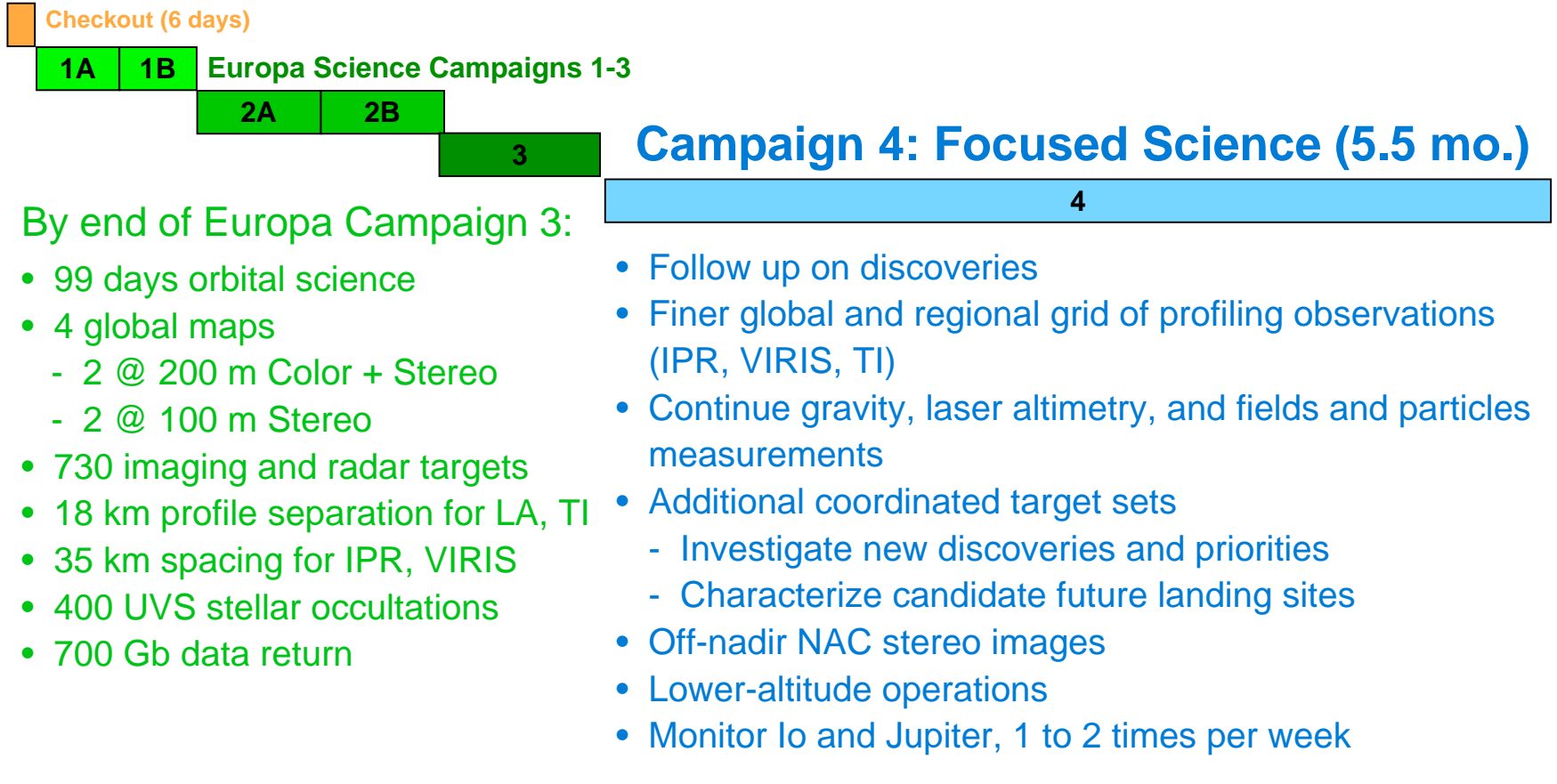


290 Mb coordinated targets

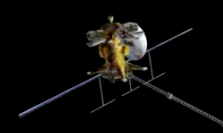
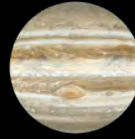
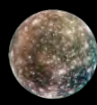
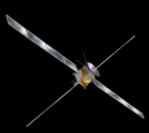
*~1700 coordinated targeted observations obtained after 9 mo.*



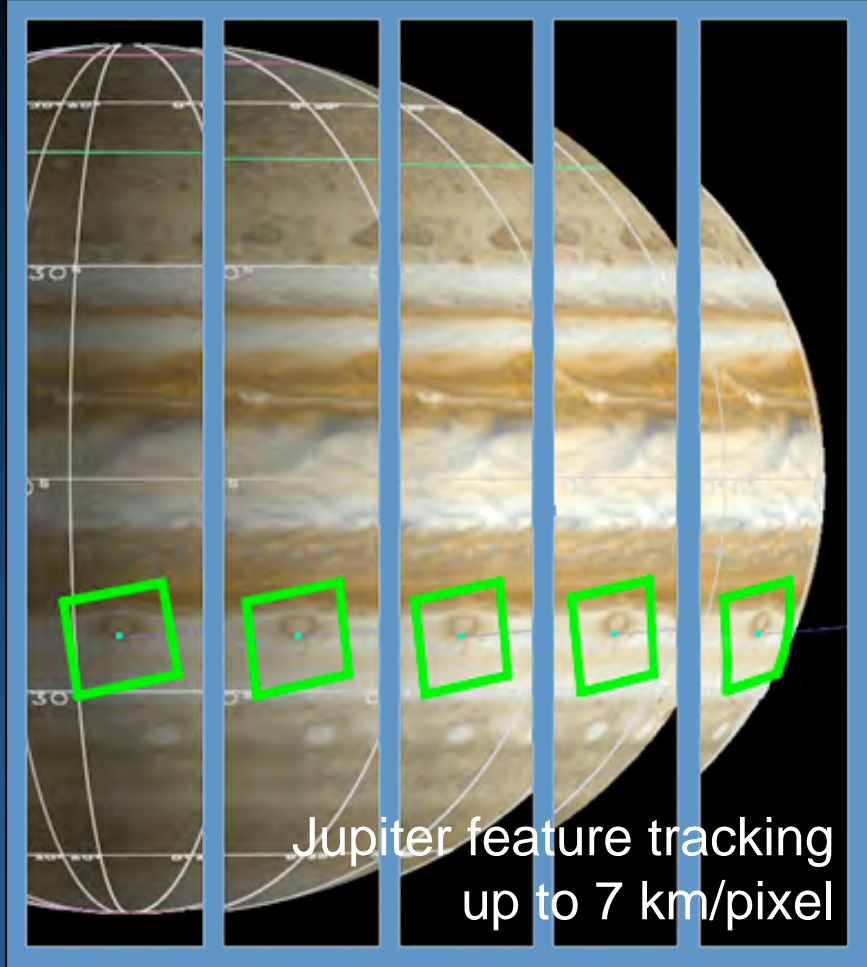
# Europa Science Campaigns



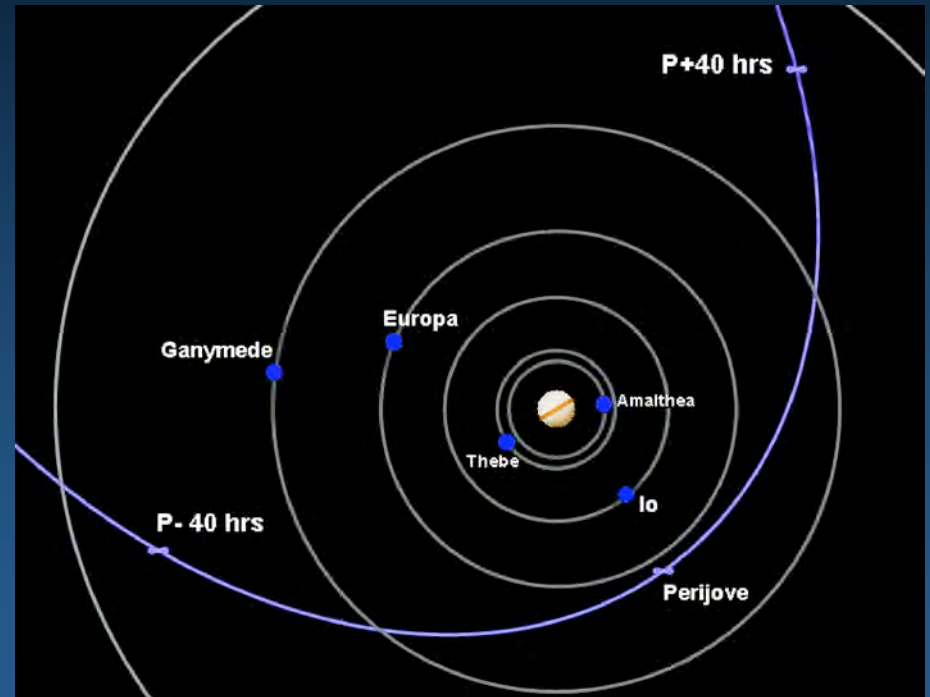
*Focused Science Campaign would allow follow-up on discoveries*



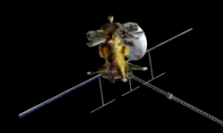
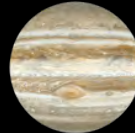
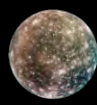
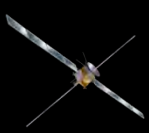
# Jupiter System Science



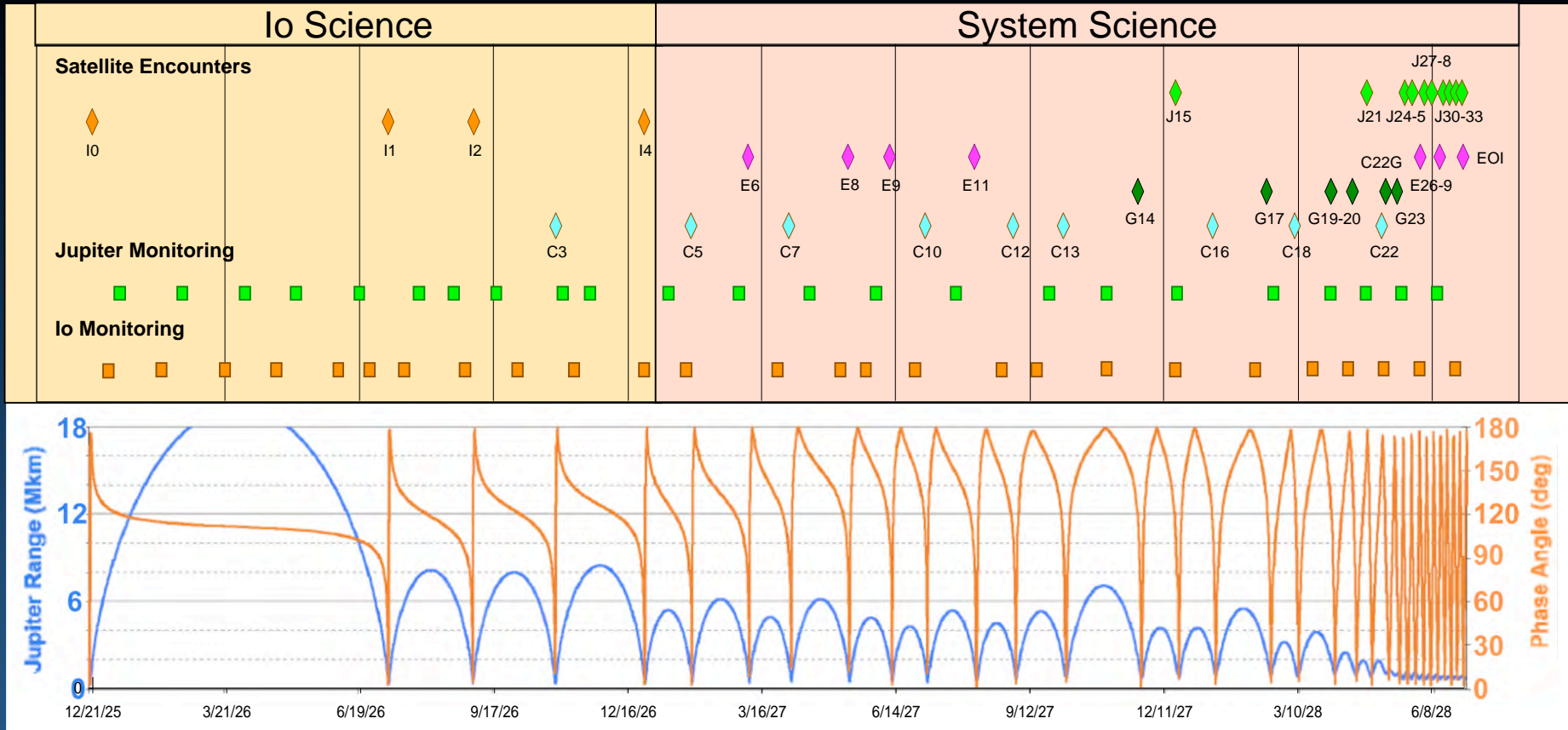
- ~25 Gb collected each perijove
- ~3.2 Tb available during Jovian tour
- ~1000 times Galileo data return



*Jovian Tour would enable in-depth Jupiter system exploration*

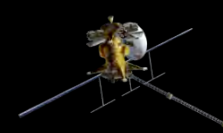
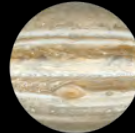
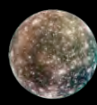
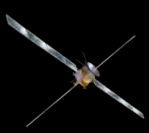


# Jovian Tour Example

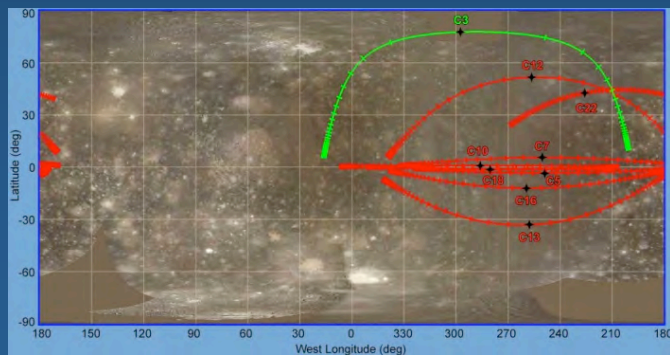
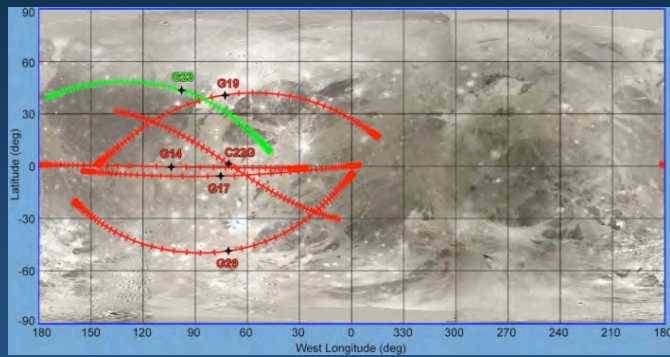
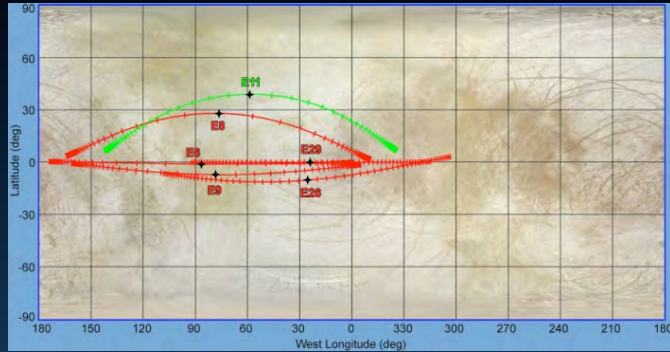


- 33 perijoves during Jovian Tour
  - 23 with satellite flybys
  - 22 permit JEO-Earth radio occultations

*Rich opportunities to acquire Jupiter System Science*

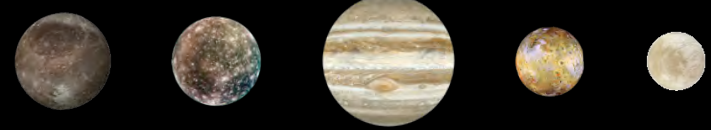
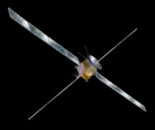


# Jovian Tour Satellite Science

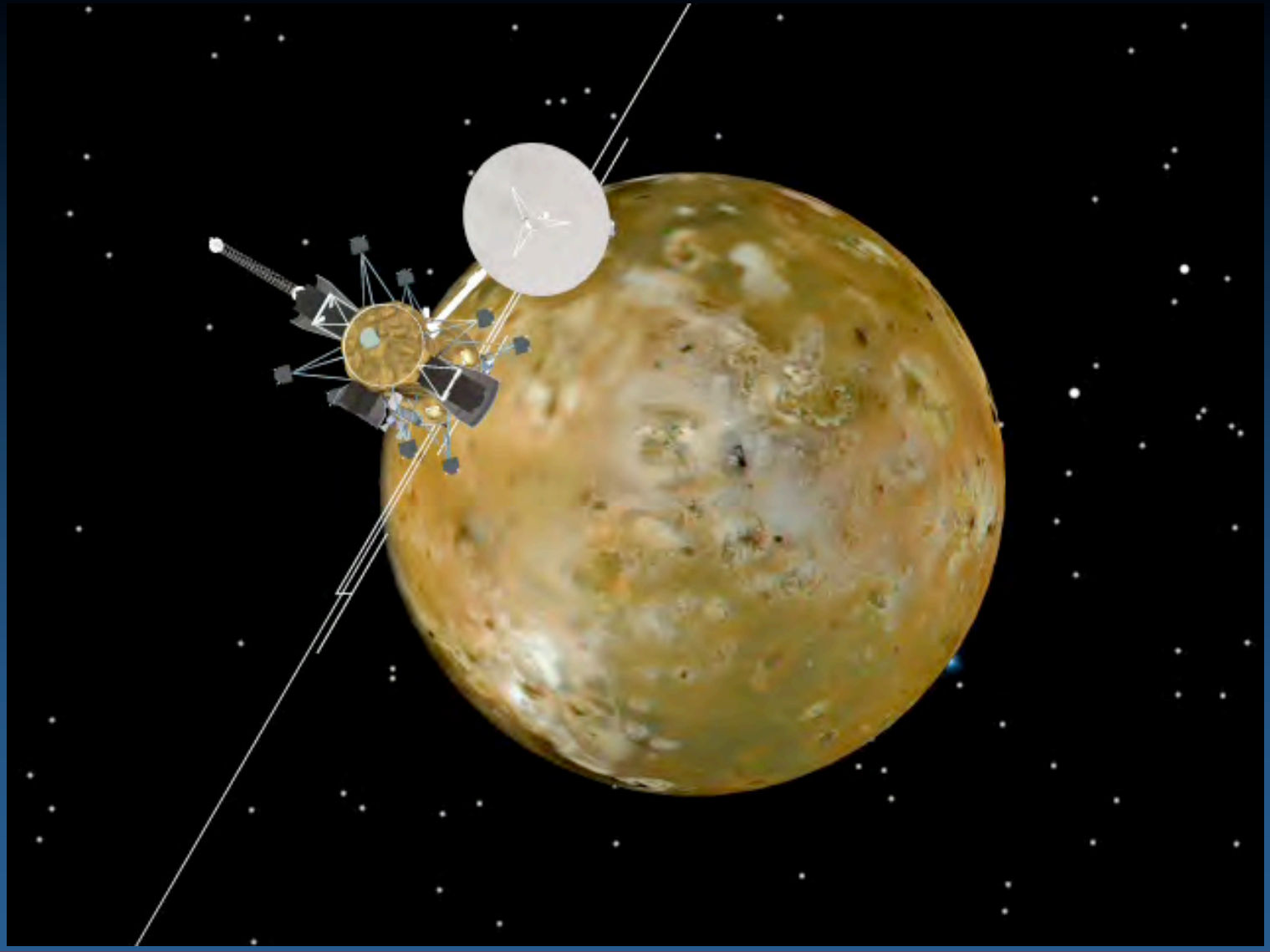


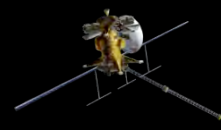
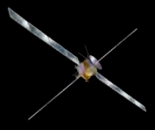
- **Io: 3 flybys**
  - Opportunities for imaging, IR spectroscopy, and altimetry
  - *In situ* analysis of extended atmosphere with INMS at 75 km
- **Europa: 6 flybys**
  - Radar and altimetry characterization and calibration
  - Imaging at up to 10–50 m resolution, NIR 250–1250 m
- **Ganymede: 6 flybys**
  - Radar sounding of grooved and dark terrains
  - Range of lats, lons for magnetosphere sampling
- **Callisto: 9 flybys**
  - High-latitude flyby for gravity field determination
  - Ocean characterization with magnetometer
  - Radar for subsurface structure of ancient cratered terrain

Satellite	≤1000m	≤200m	≤50m	≤10m	Length IPR (km)	Length LA (km)
<b>Io</b>	30%	20%	5%	-	1000	7400
<b>Europa</b>	60%	60%	15%	0.01%	6600	19000
<b>Ganymede</b>	50%	50%	10%	0.02%	17000	28000
<b>Callisto</b>	85%	75%	5%	0.01%	15000	30000



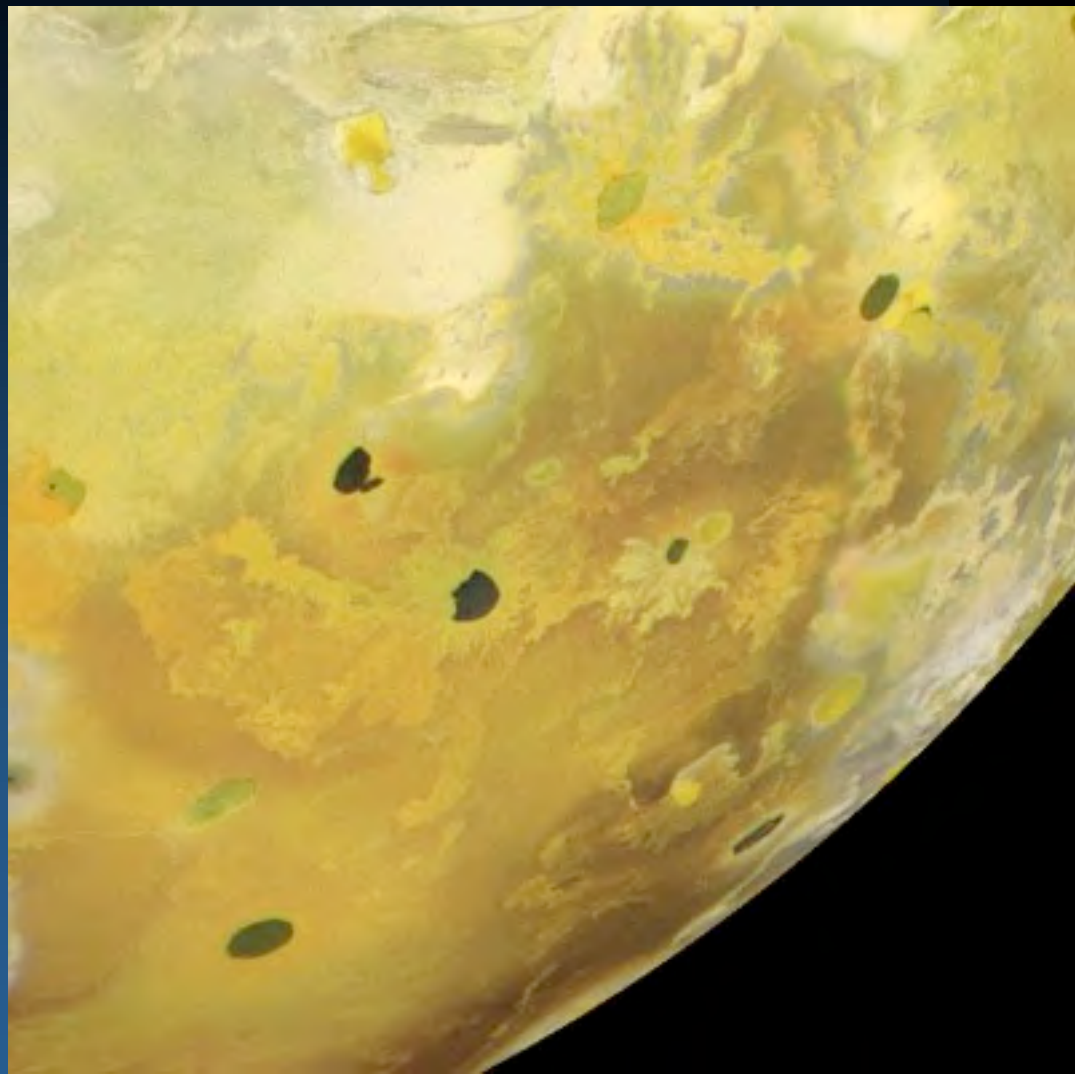
# Io Flyby Example



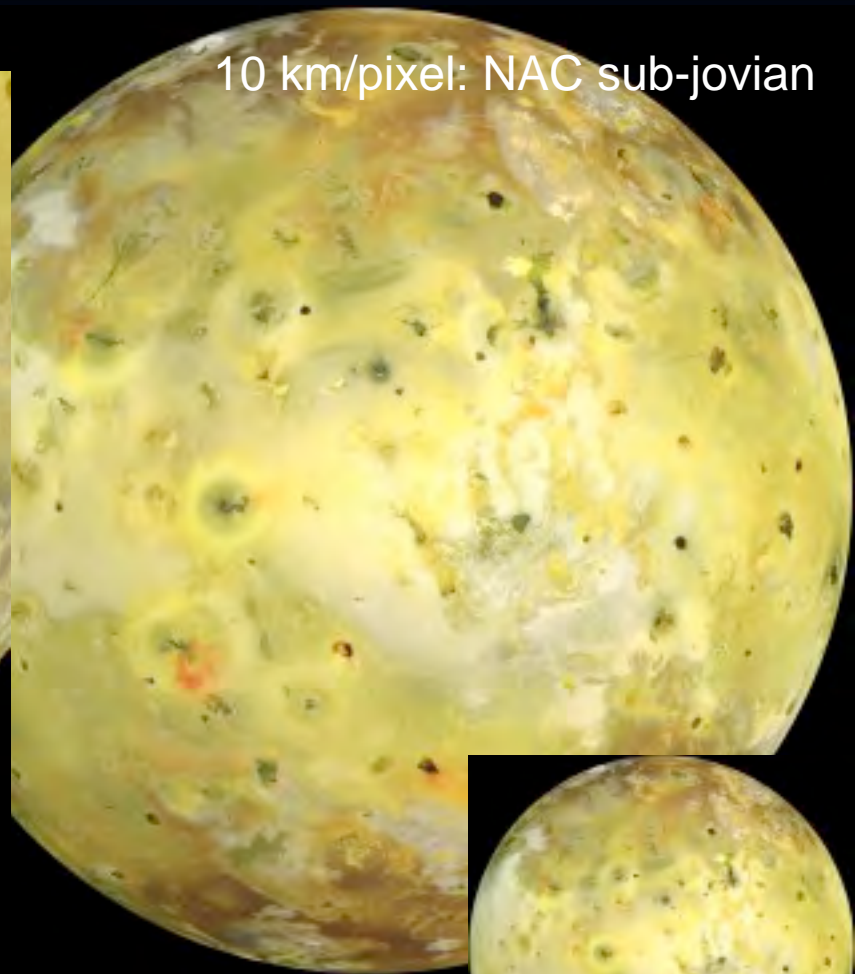


# Io Resolution Examples from Europa Distance

2.5 km/pixel: NAC anti-jovian

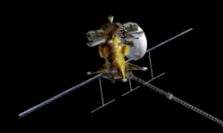
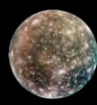
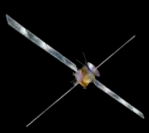


10 km/pixel: NAC sub-jovian

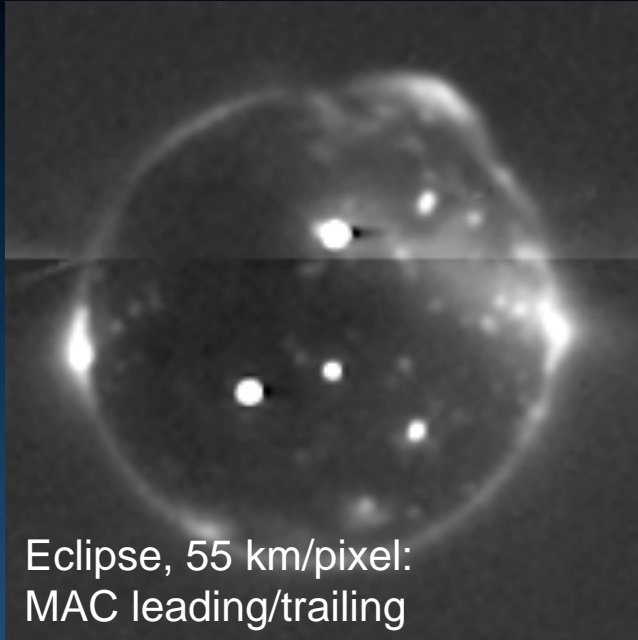


25 km/pixel: MAC anti-jovian





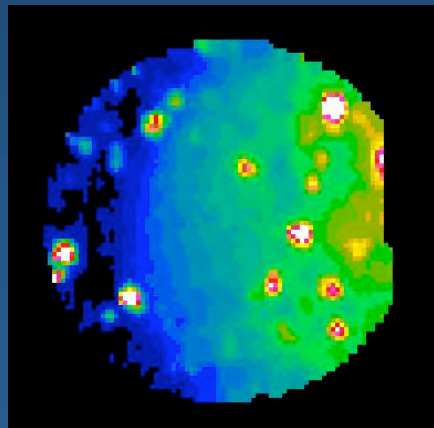
# Io Resolution Examples from Europa (cont.)



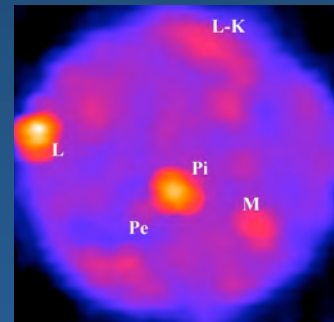
Eclipse, 55 km/pixel:  
MAC leading/trailing



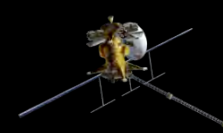
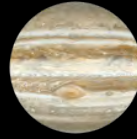
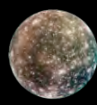
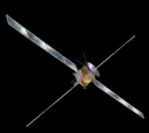
High phase, 25 km/pixel:  
MAC anti-jovian



Near IR (5  $\mu\text{m}$ ),  
85 km/pixel:  
VIRIS ~anti-jovian

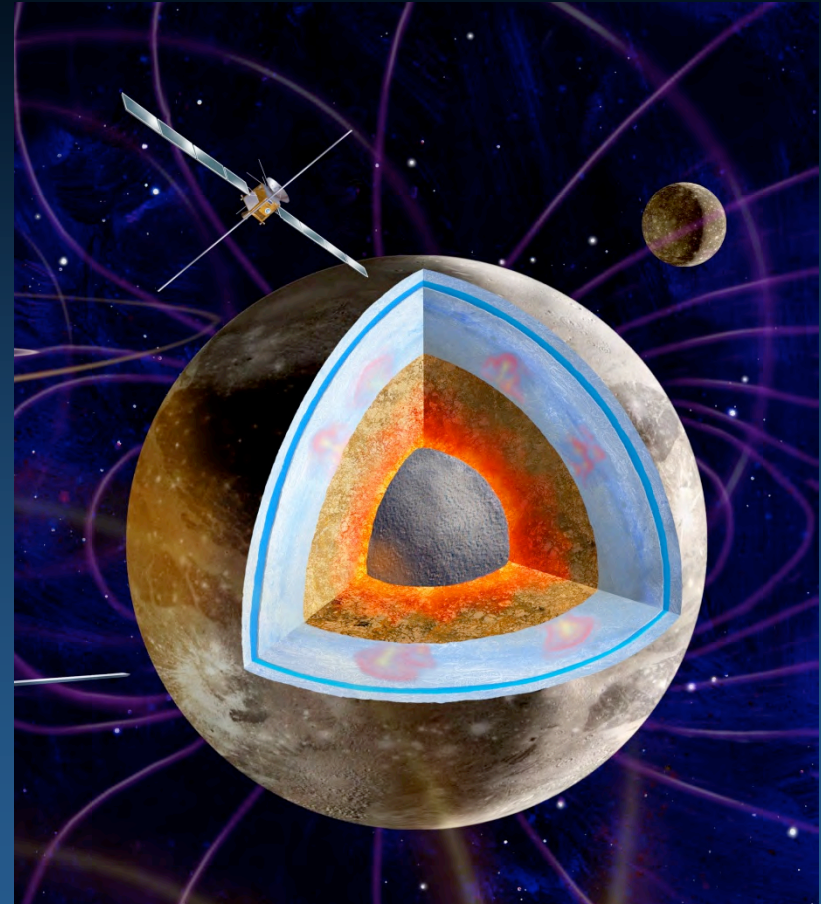


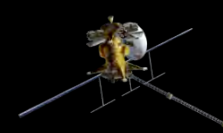
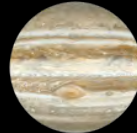
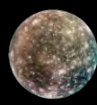
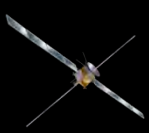
Thermal IR  
(~30  $\mu\text{m}$ ),  
350 km/pixel:  
anti-jovian  
[Twice as good as  
JEO thermal mapper]



# JGO Science: Overview

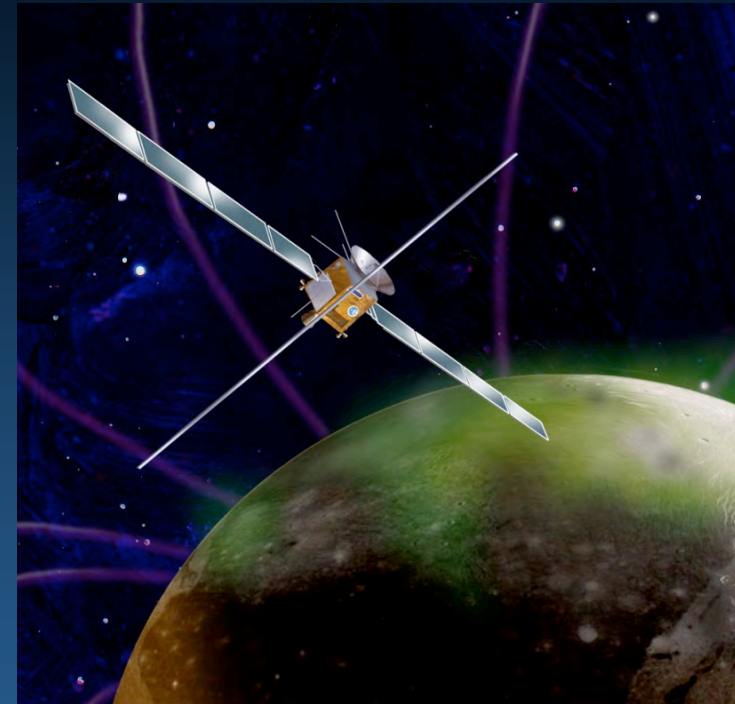
- Key JGO science phases
  - **Jupiter system:** In-depth exploration
    - From Jupiter orbit, synergistically with JEO
  - **Callisto:** In-depth study and mapping
    - Multiple flybys using a resonant orbit
  - **Ganymede:** Detailed orbital study
    - Elliptical orbit first, then circular orbit
  
- Science Objectives:
  - **Ganymede:** Characterize Ganymede as a planetary object, including its potential habitability
  - **Satellite System:** Study the Jovian satellite system
  - **Jupiter:** Study the Jovian atmosphere
  - **Magnetosphere:** Study the Jovian magnetodisk / magnetosphere
  - **Jupiter system:** Study the interactions occurring in the Jovian system

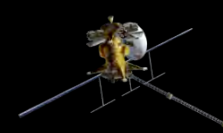
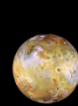
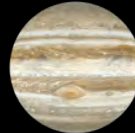
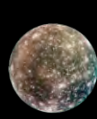
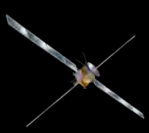




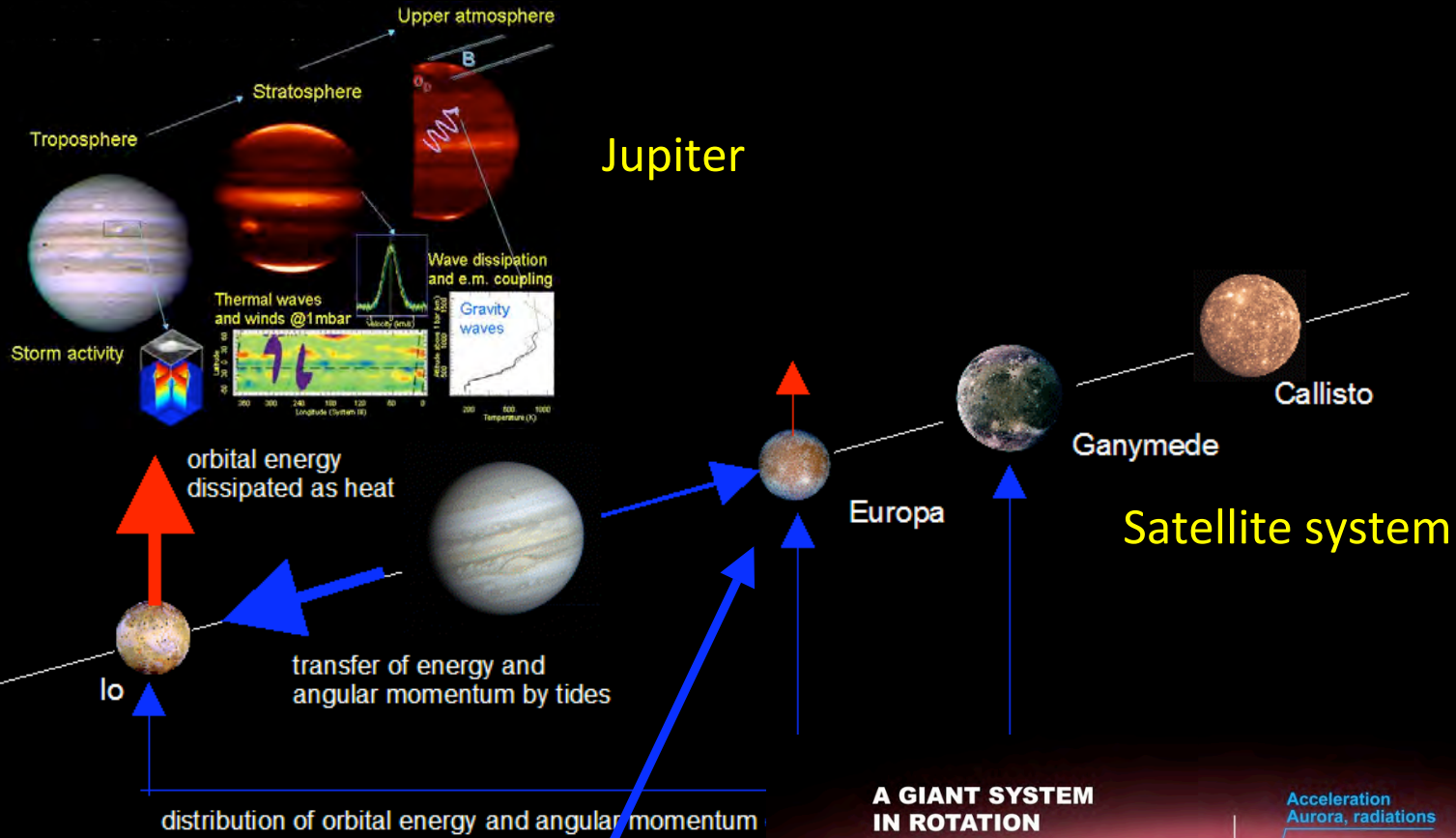
# JGO Payload Definition Document (PDD) Study Model Payload

PDD Model Instrument Name	Acronym
Medium-Res Camera & Wide Angle Camera	WAC+MRC
Magnetometer	MAG
Radio Science Transponder and USO	JRST+USO
Visible InfraRed Hyperspectral Imaging Spectrometer	VIRHIS
Plasma Package & Ion and Neutral Mass Spectrometer	PLP/INMS
<b>Sub-mm Instrument</b>	<b>SWI</b>
<b>Radio and Plasma Wave Instrument</b>	<b>RPWI</b>
Narrow Angle Camera	HRC
Sub-Surface Radar	SSR
Laser Altimeter	LA
UV Imaging Spectrometer	UVIS

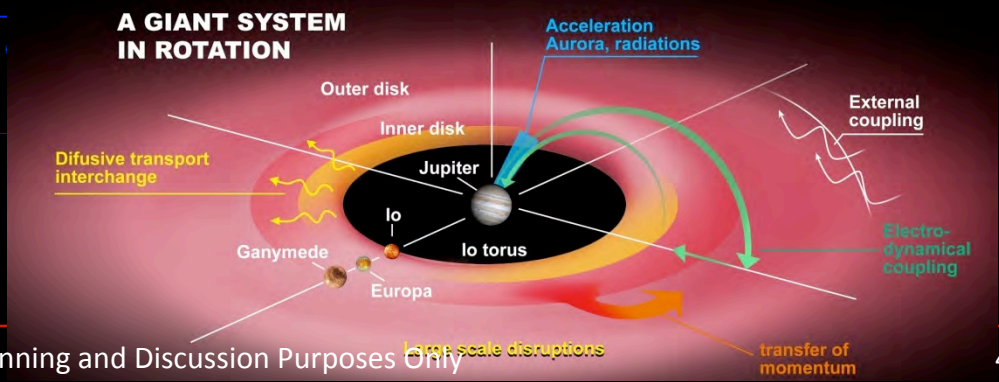


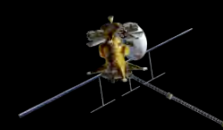
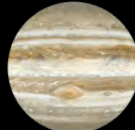
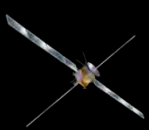


# The Jupiter System: Three Coupled Components



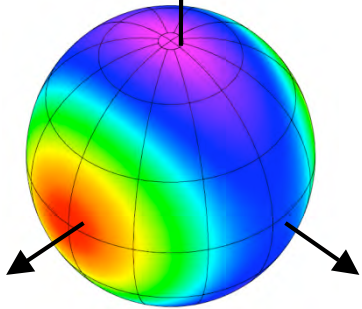
**Magnetodisk/  
radiation belts**





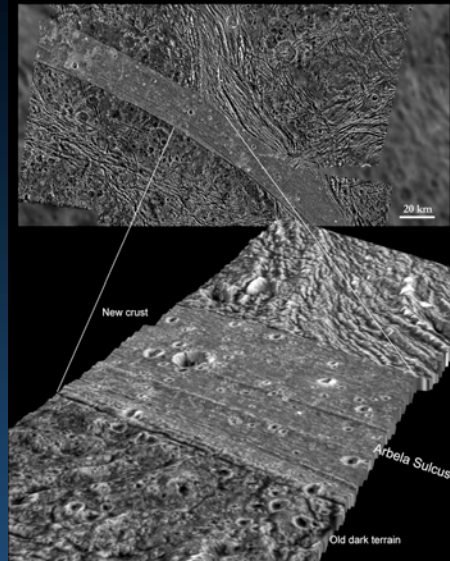
# Ganymede: Europa's "False Twin"

Tidal deformation

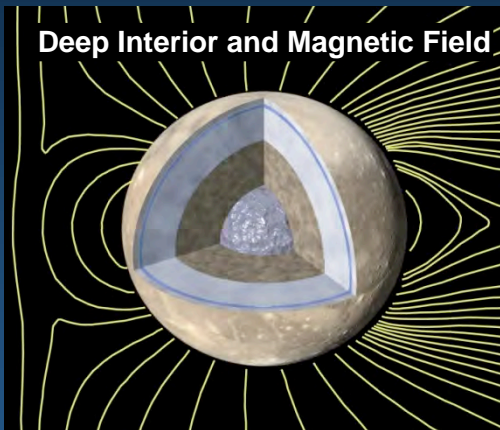


- Presence and extent of a subsurface ocean
- Ice shell and subsurface water
- Deep internal structure, dynamo, magnetic field
- Coupling among surface, exosphere, and magnetosphere
- Surface composition and chemistry
- Surface features, tectonic processes
- Thermal evolution, geology, and the Laplace resonance

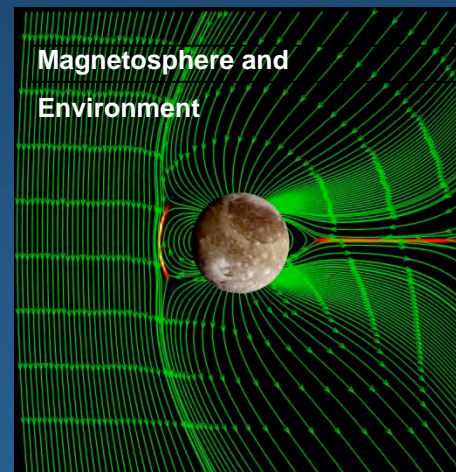
Geology and Topography



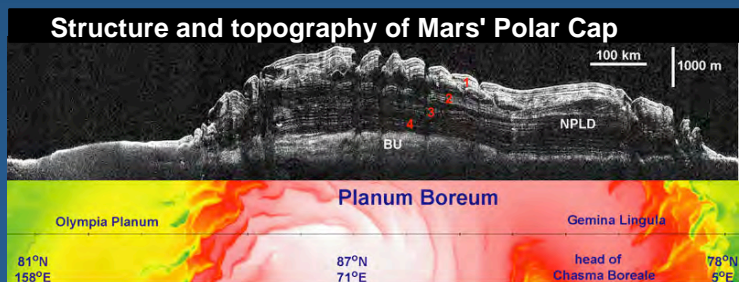
Deep Interior and Magnetic Field



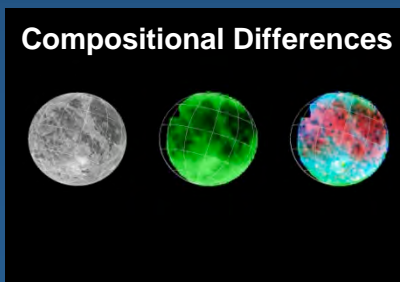
Magnetosphere and Environment

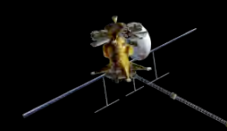
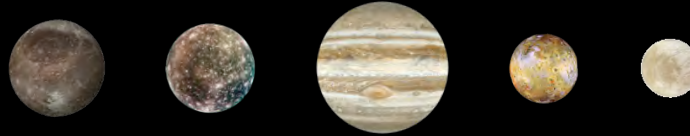
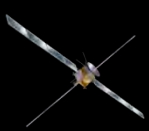


Structure and topography of Mars' Polar Cap



Compositional Differences

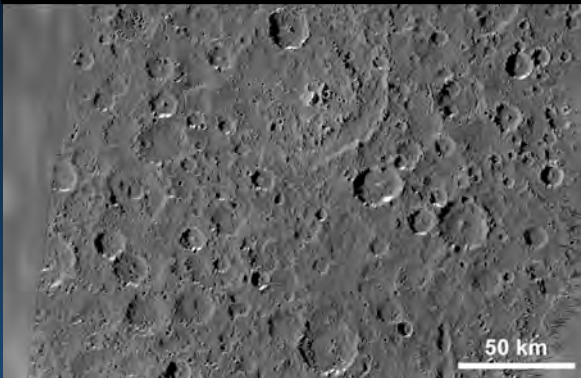




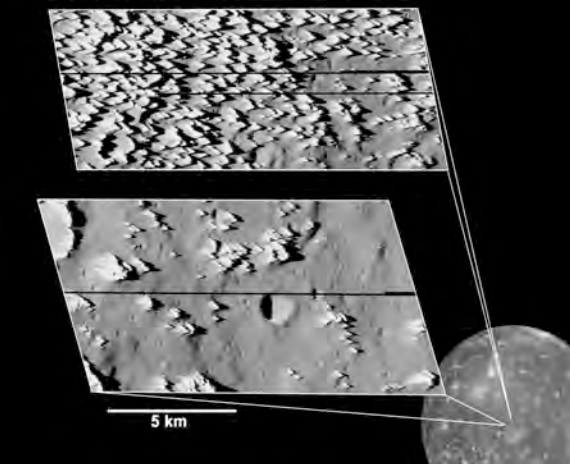
# Callisto: A Witness of the Early Ages

- Presence and extent of a subsurface ocean
- Ice shell and subsurface water
- Deep internal structure, including degree of differentiation
- Cratering record and early geological history
- Surface composition, including organics and CO<sub>2</sub>
- Surface degradational processes (erosion and sublimation)

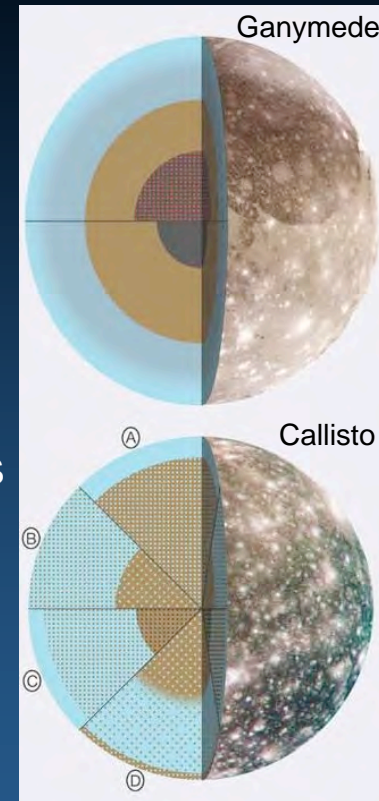
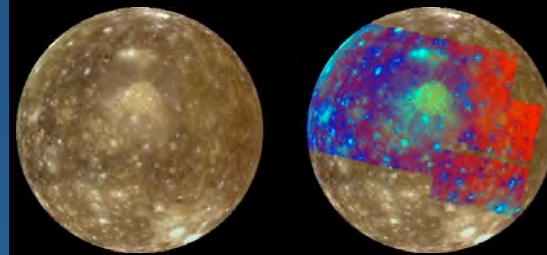
Crater Distribution and Morphology



Knobby Terrain: Erosion Processes

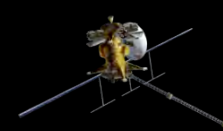
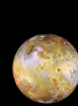
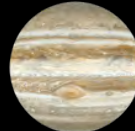
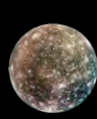
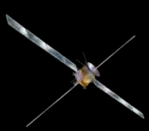


Compositional Heterogeneities

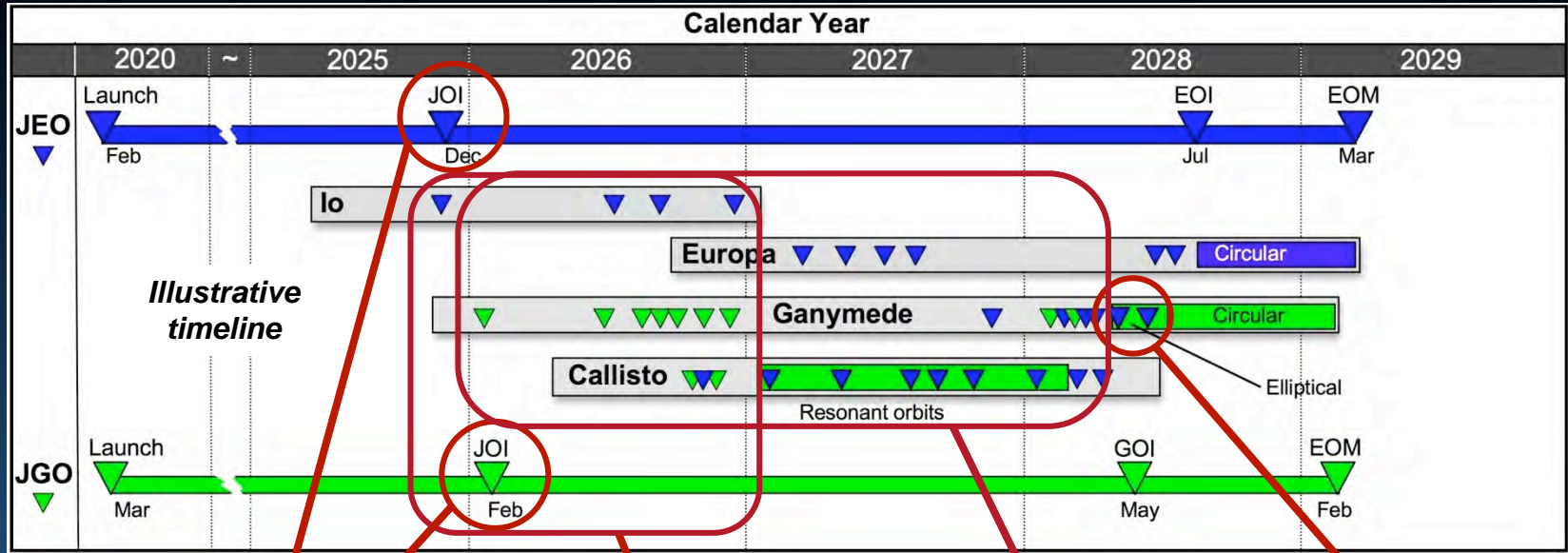


Internal differentiation:  
Where is Callisto ?

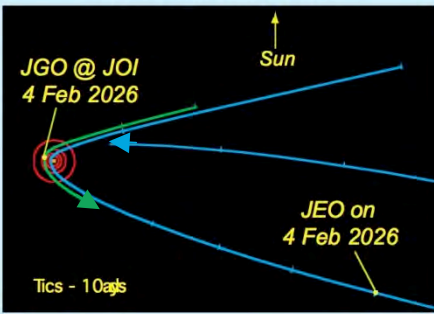
Image after Bagenal et al. [2004]



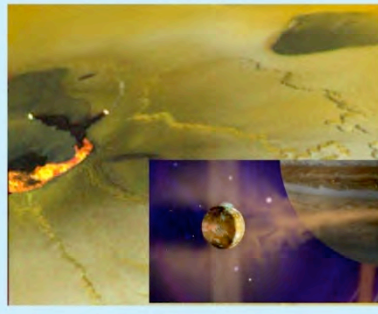
# EJSM Synergistic Science



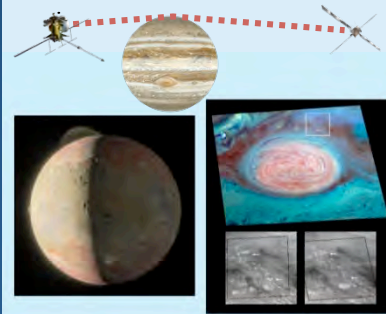
## Jupiter Magnetosphere Studies



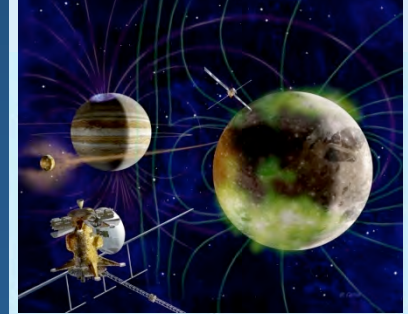
## Io Volcanism & Io Torus Dynamics

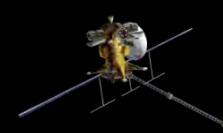
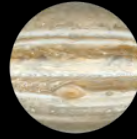
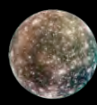
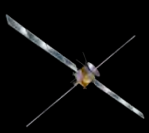


## Satellite & Jupiter Monitoring; Radio Occultation Science?



## Ganymede Magnetosphere Studies



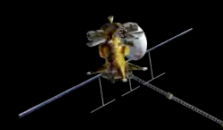
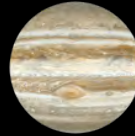
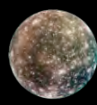
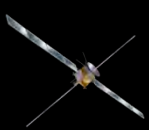


# JEO: Untangling Fundamental Mysteries

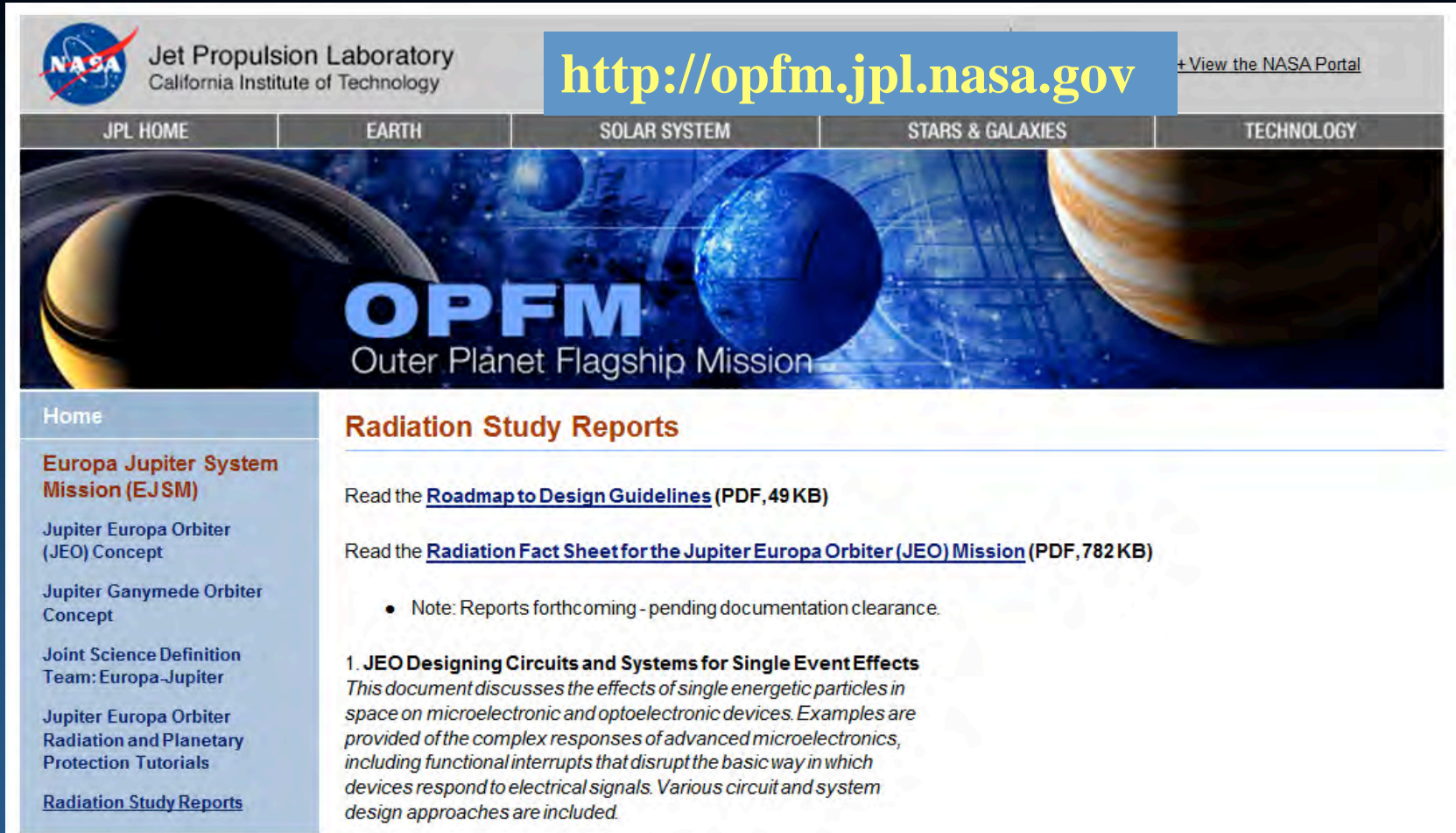
- Science objectives have been honed for over a decade
- Science and technical teams have worked together closely to optimize the JEO science return
- JEO would achieve outstanding Decadal Survey science
- JEO would fundamentally alter our understanding of the nature of habitable worlds around gas giants
- EJSM has the science, technology, and “opportunity”



*Solving fundamental mysteries of Europa and the Jupiter system*



# For More Information...



Jet Propulsion Laboratory  
California Institute of Technology

<http://opfm.jpl.nasa.gov> [View the NASA Portal](#)

JPL HOME EARTH SOLAR SYSTEM STARS & GALAXIES TECHNOLOGY

## OPFM

Outer Planet Flagship Mission

Home

- Europa Jupiter System Mission (EJSM)**
  - Jupiter Europa Orbiter (JEO) Concept
  - Jupiter Ganymede Orbiter Concept
  - Joint Science Definition Team: Europa-Jupiter
  - Jupiter Europa Orbiter Radiation and Planetary Protection Tutorials
  - [Radiation Study Reports](#)

### Radiation Study Reports

Read the [Roadmap to Design Guidelines](#) (PDF, 49 KB)

Read the [Radiation Fact Sheet for the Jupiter Europa Orbiter \(JEO\) Mission](#) (PDF, 782 KB)

- Note: Reports forthcoming - pending documentation clearance.

**1. JEO Designing Circuits and Systems for Single Event Effects**  
*This document discusses the effects of single energetic particles in space on microelectronic and optoelectronic devices. Examples are provided of the complex responses of advanced microelectronics, including functional interrupts that disrupt the basic way in which devices respond to electrical signals. Various circuit and system design approaches are included.*

*A wealth of resources are available at the OPFM website*

# The Europa Jupiter System Mission: The Emergence of Habitable Worlds Around Gas Giants

